

Using 2DFFT method to search Fast Radio Bursts

牛晨辉 (Chenhui Niu)

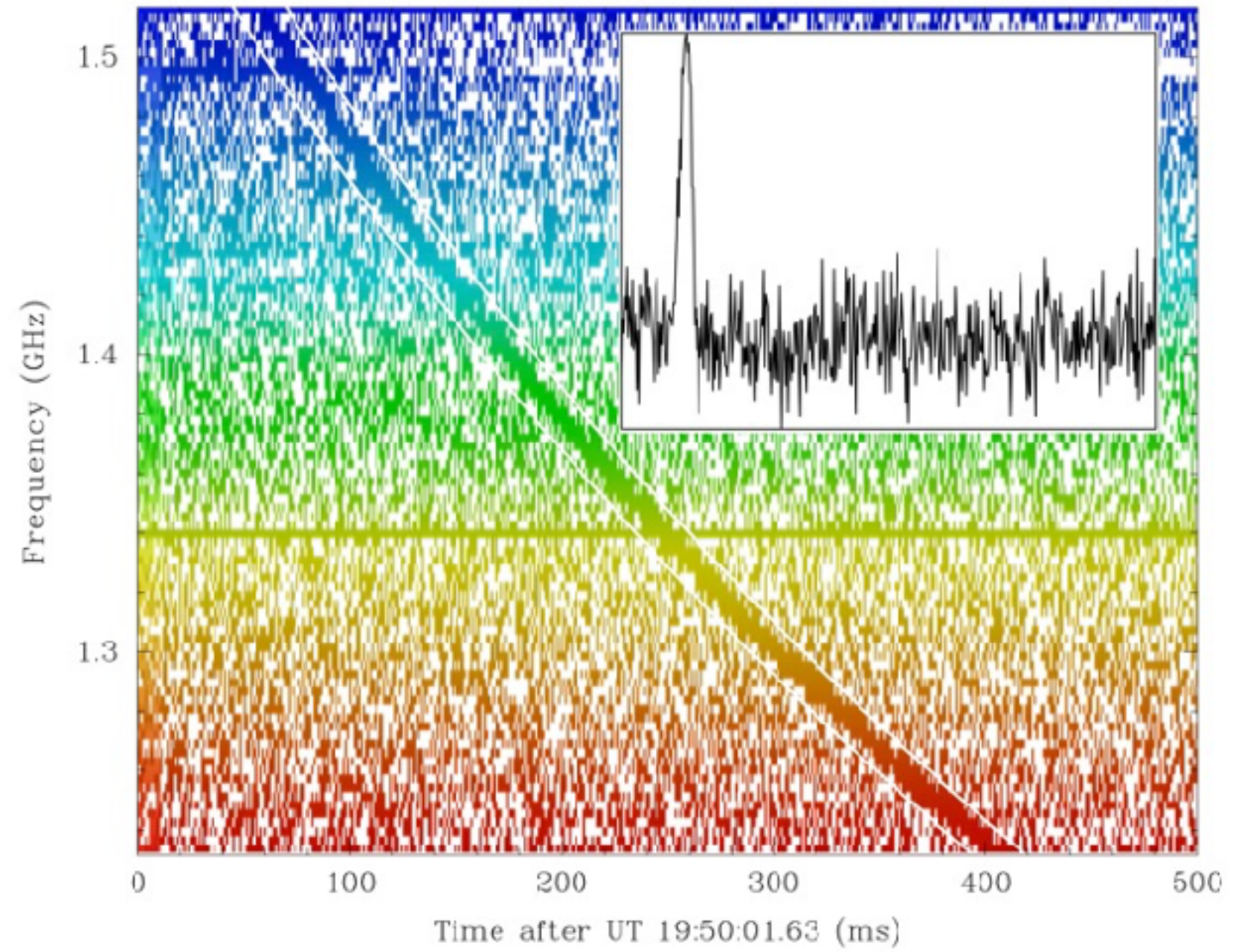
华中师范大学 (Central China Normal University)

Collaborator : Yichao Li, Xuelei Chen^{*}, Ue-li Pen, Fengquan Wu et al.

2023 July 19th
Shenyang

Lorimer Burst

Lorimer Burst



(Lorimer et al. 2007)

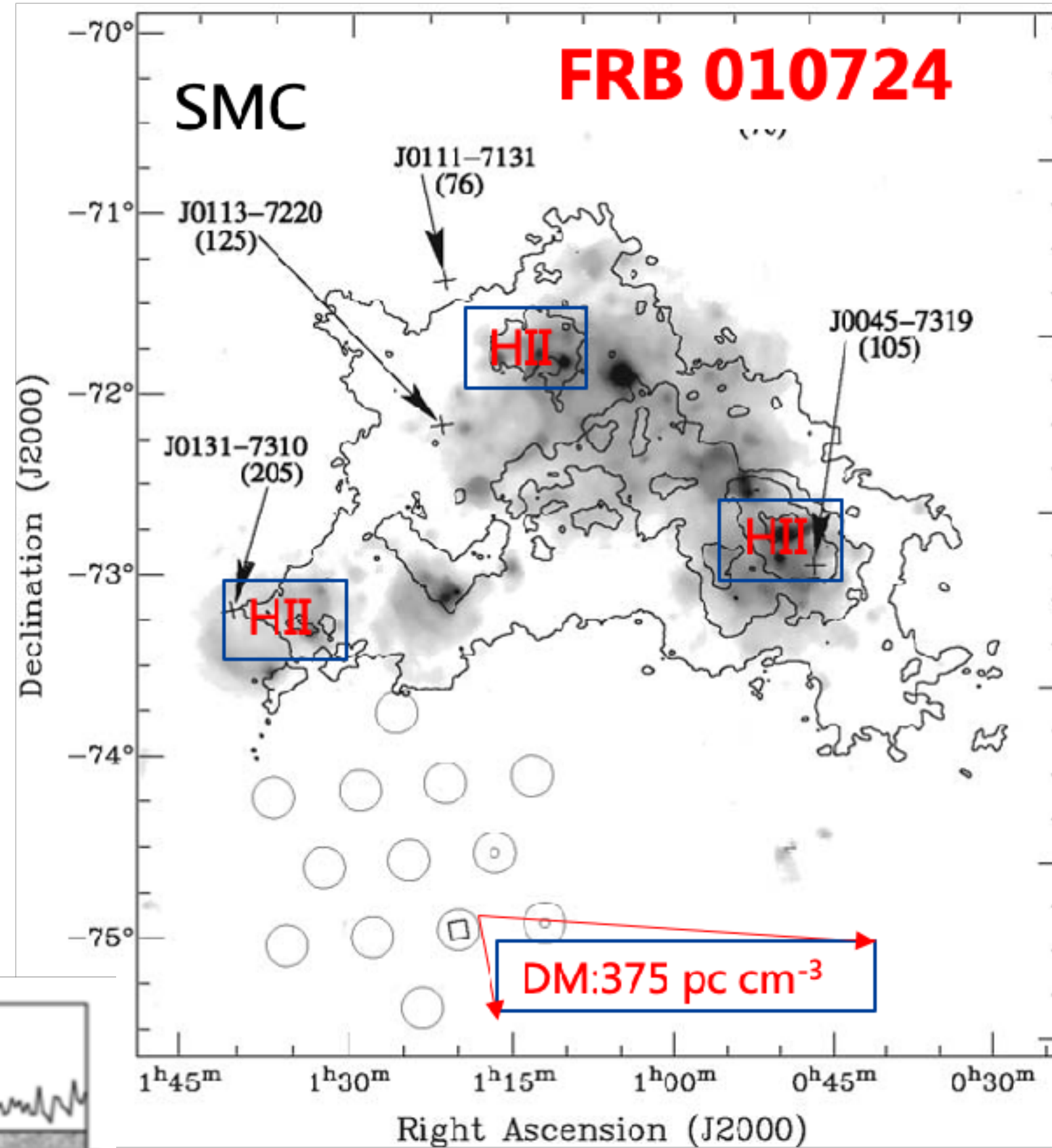
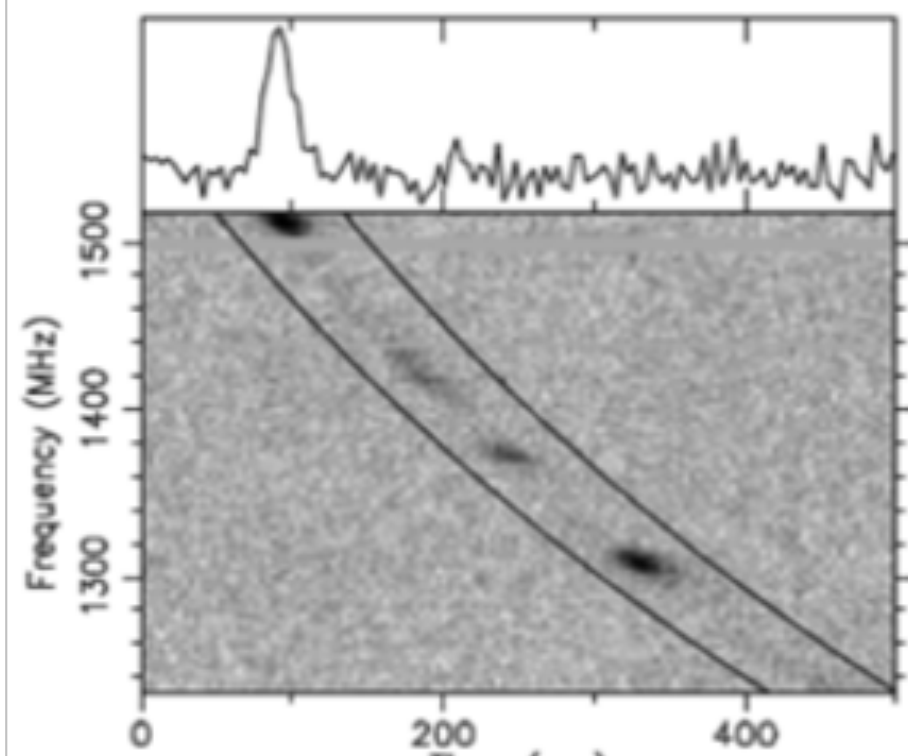
High DM Pulsar ?

No follow-up detection:

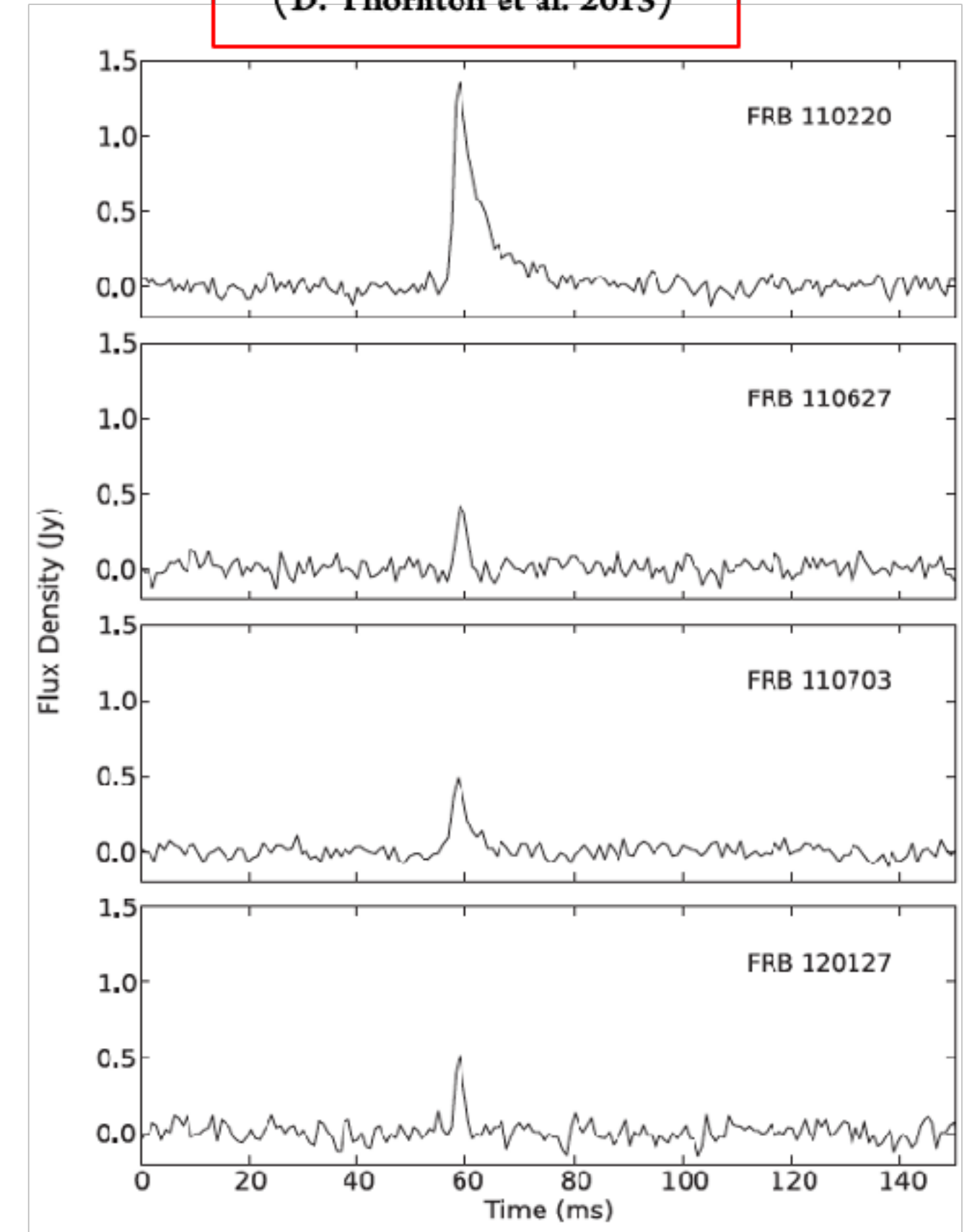
- 50 hours 2001
- 40 hours 2007
- Trail DM : 0-500 pc cm⁻³
- High Galactic latitude

Peryton?

- Only show in 1/13 beam



(D. Thornton et al. 2013)

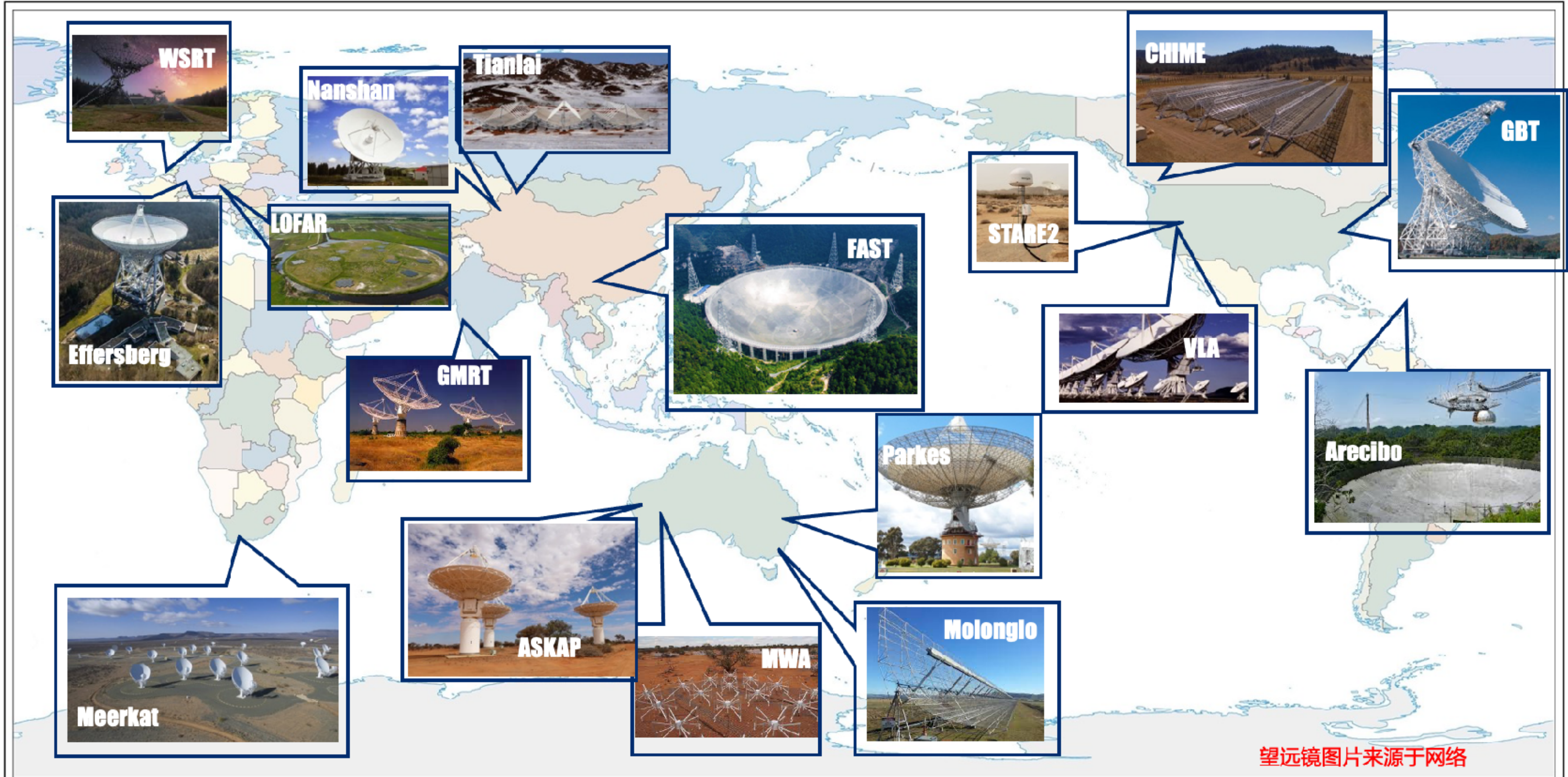


$DM_{MK} < 6\% DM_{Obs}$

- High DM
- High Galactic latitude
- Duration: ~ms

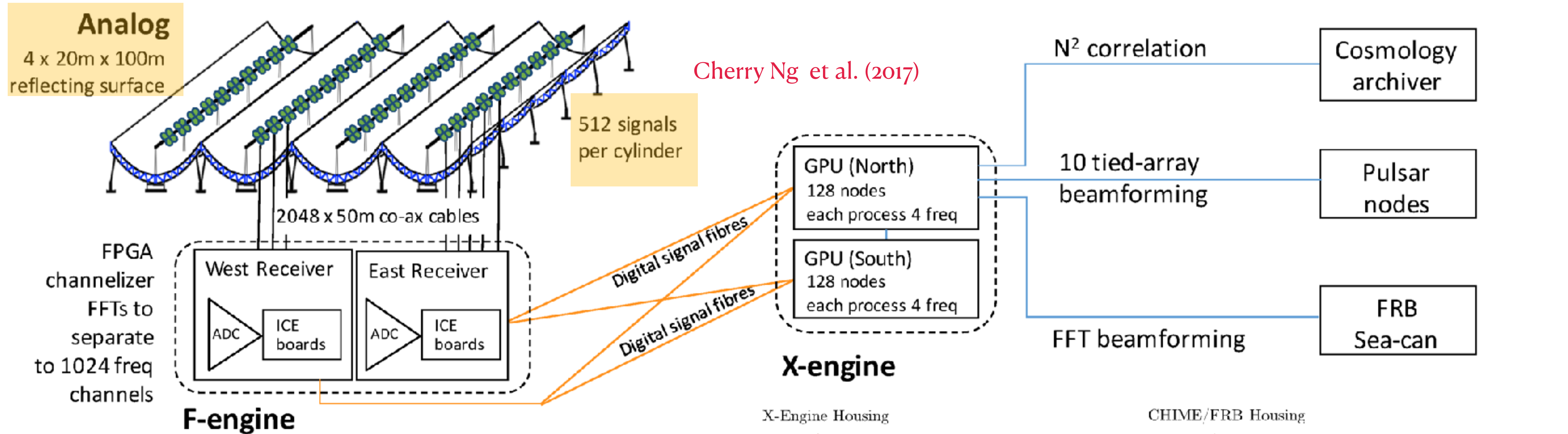
The FRB Surveys

世界地图

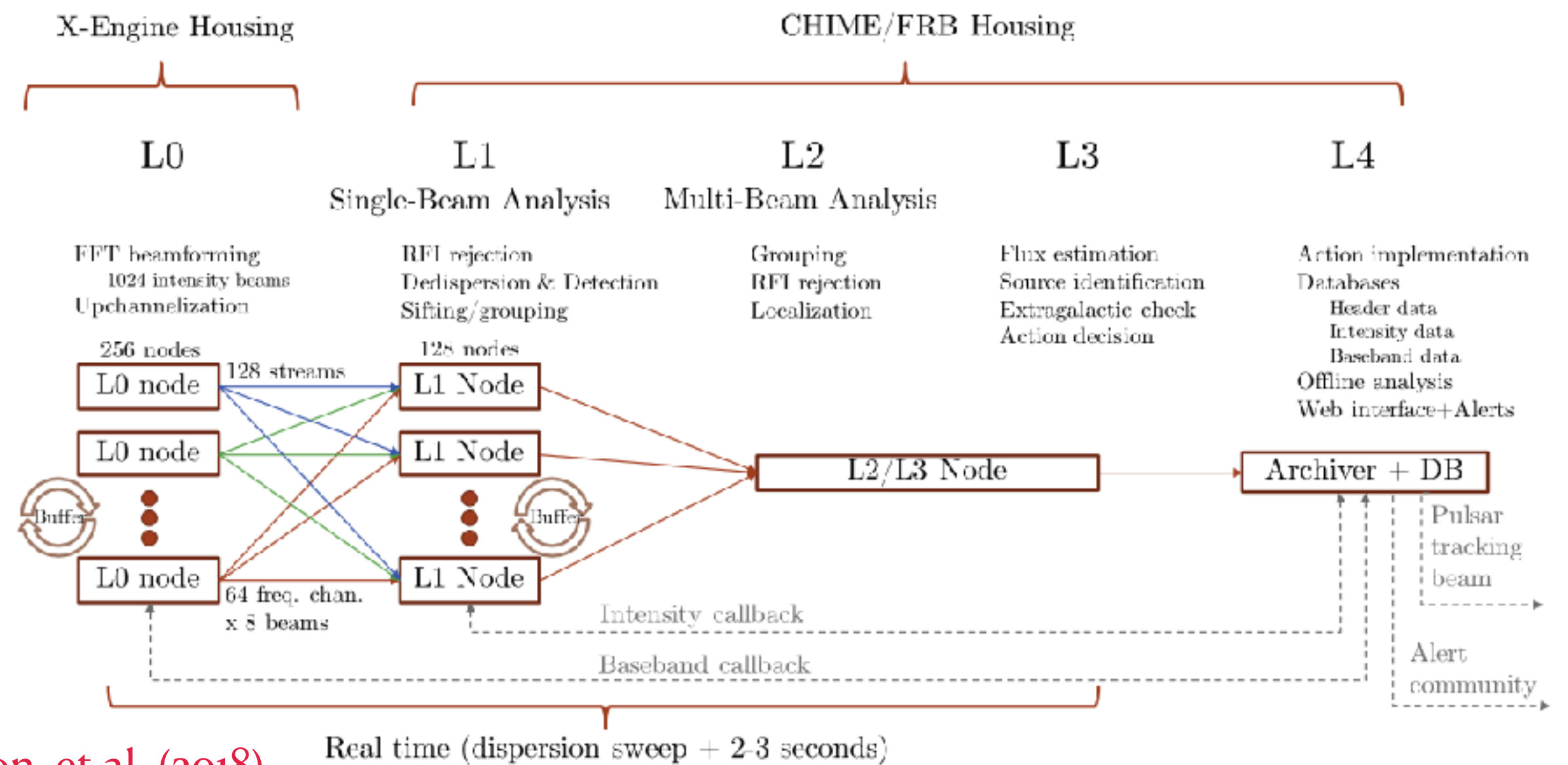


望远镜图片来源于网络

CHIME



Cherry Ng et al. (2017)



CHIME Collaboration et al. (2018)

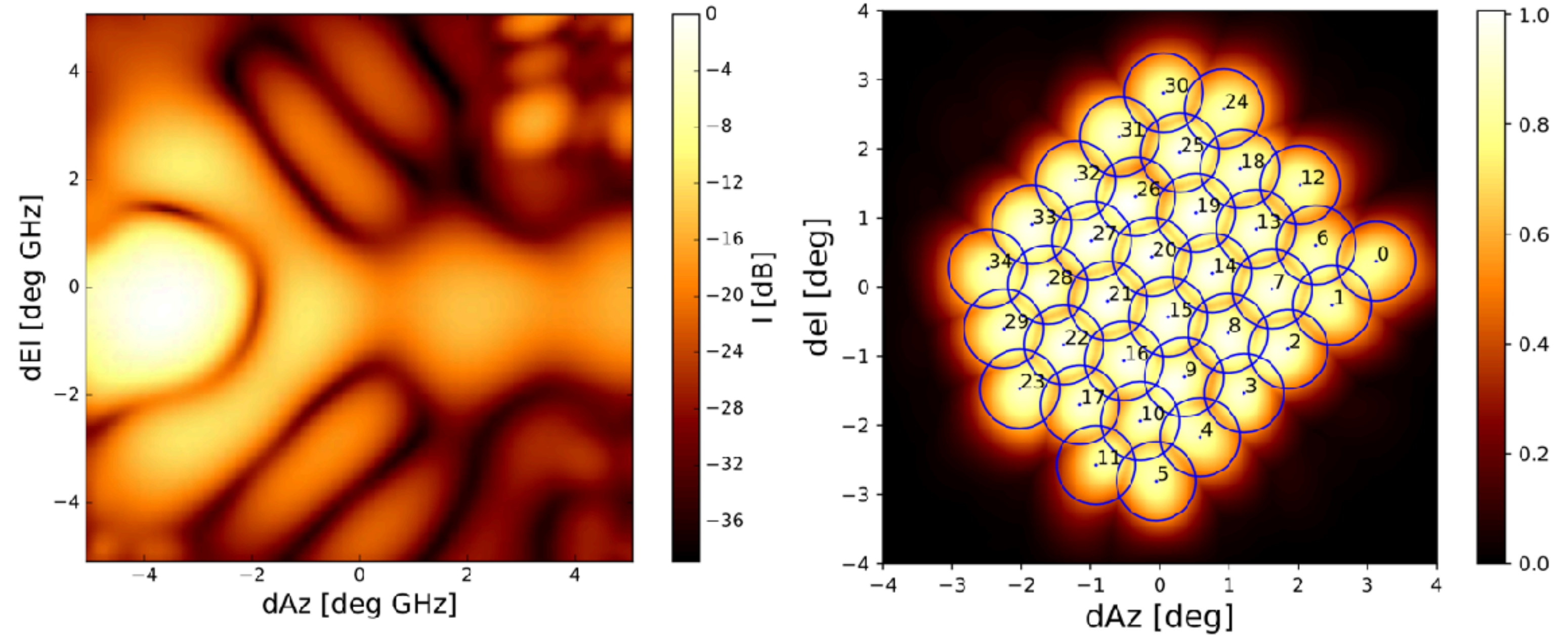
ASKAP

C. W. James, et al, 2019

Phase Array Feed



Dimeter :12 m
Beam: ~1 deg
PAF: ~5.5 deg



Phase Array Feed

Fly's eye mode

- Use 36 beams in each path
- Use 36 antennas pointing independently
- Effectively $36 \times 36 =$ beams
- Powerful tool for finding fast transients

MeerKat

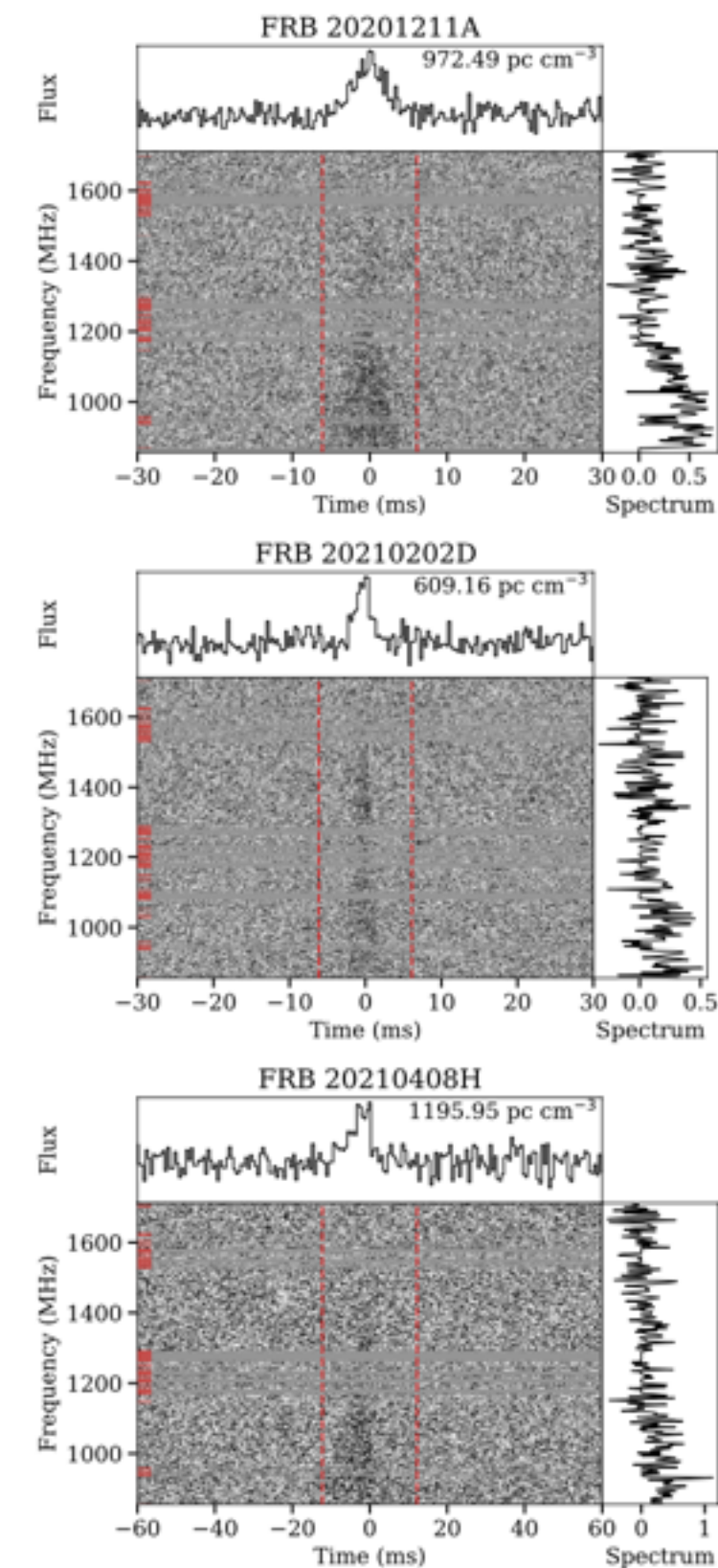


- 64 13.5-m telescopes
- Central compact array
- South Africa
- Along with the HERA, ASKAP and the MWA, precursors to the final SKA.

Specifications

Number of antennae	64
Dish diameter	13.5 m
Minimum baseline	29 m
Maximum baseline	8 km
Frequency bands (receivers)	0.58 – 1.015 GHz 0.9 - 1.67 GHz 8 – 14.5 GHz
Continuum imaging dynamic range at 1.4 GHz	60 dB
Line-to-line dynamic range at 1.4 GHz	40 dB
Mosaicing imaging dynamic range at 1.4 GHz	27 dB
Linear polarisation cross coupling across -3 dB beam	-30 dB

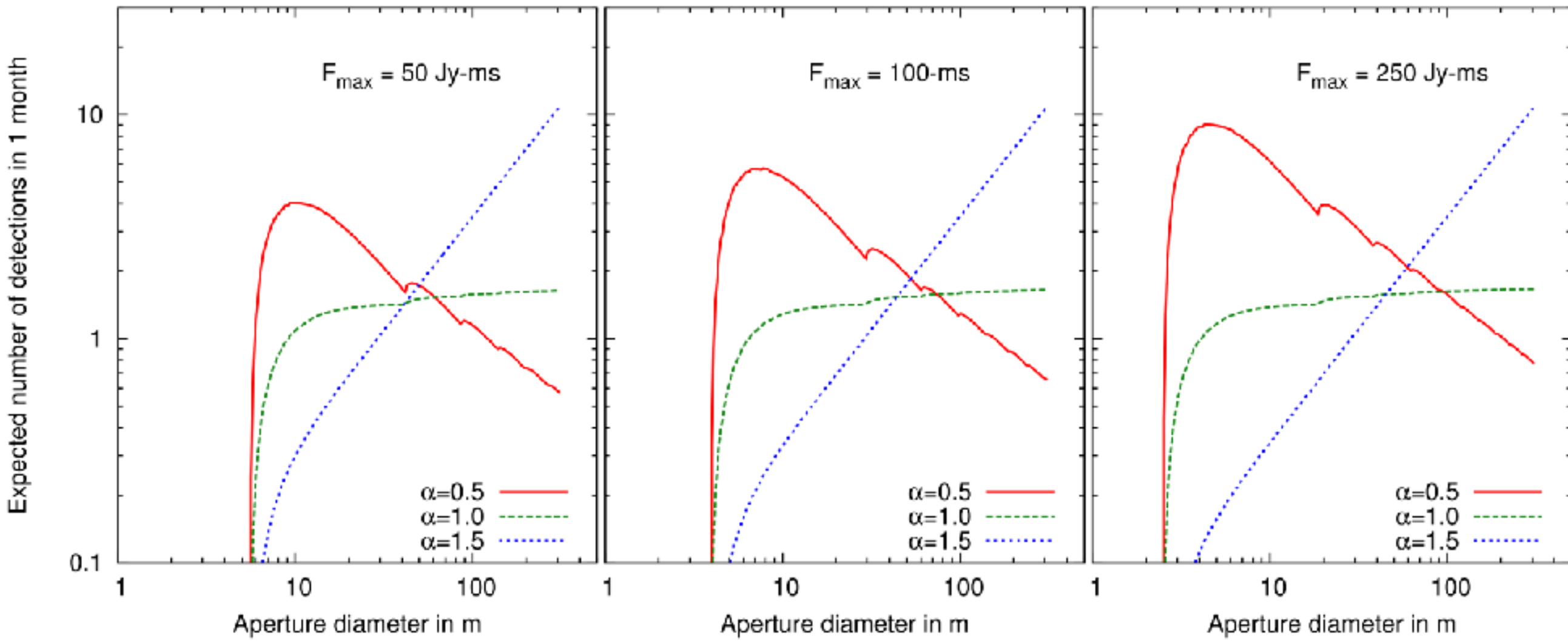
The system overview is on going...



Deep Synoptic Array

Vedantham et al. (2016)

Condition: $T_{sys}=60$ K, $\eta=60\%$, $\Delta \nu = 500$ MHz, $\nu_c=1.5$ GHz, $N_{stations} = 10$



The relatively shallow slope of the FRB logN–logF curve implies that a modest telescope array with $N \sim 10$ and $d < 6$ m aperture is sufficient to detect and localize a large population (> 1 month $^{-1}$) of FRBs.

- 110 × 4.65-m dishes
- Owens Valley Radio Observatories
- continuously survey for FRBs at 1280 – 1530 MHz
- Localization < 3"
- 100 FRBs per year

Parameter	Value
Number of dishes *	110
Dish diameter	4.65 m
Aperture efficiency	0.65
System temperature †	25 K
Frequency band	1280 – 1530 MHz
Number of channels	8192
FRB search sensitivity (7.5σ , 1 ms) ‡	< 1.7 Jy ms
Field of view	10 deg ²
Maximum baseline	2.6 km
FRB localization accuracy	± 1.5 arcsec
Survey speed	3.1×10^4 deg ² m ⁴ K ⁻²



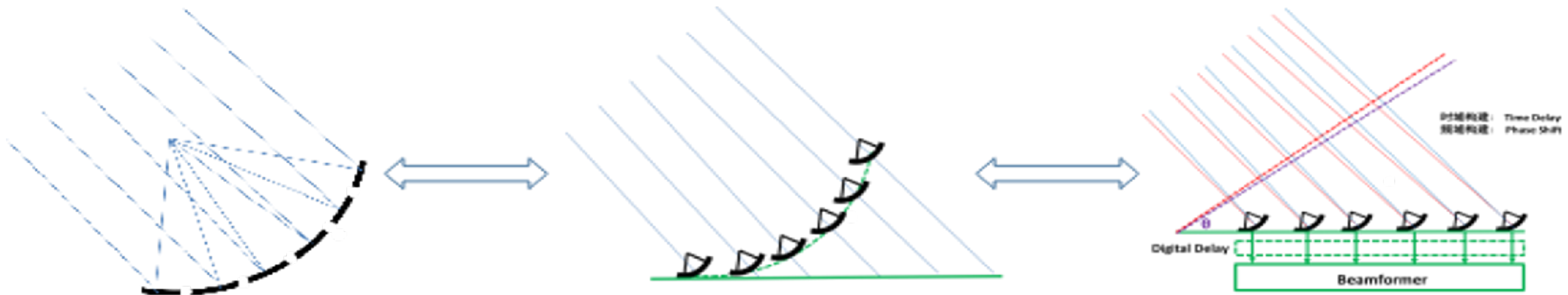
Small Dishes Interferometry

Advantage:

- Portable
- Matured technique
- Extensible deployment
- Low Coast

Disadvantage:

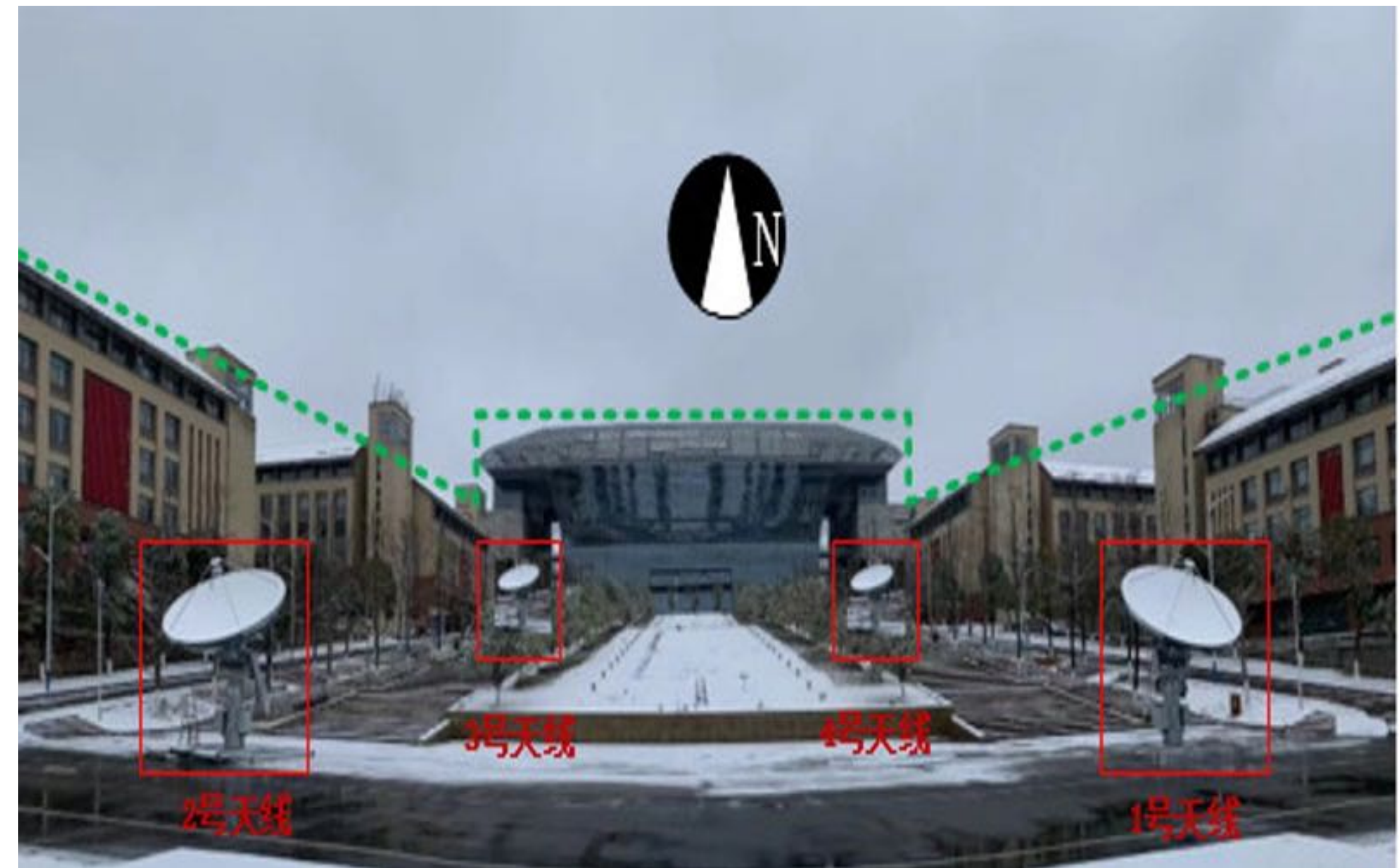
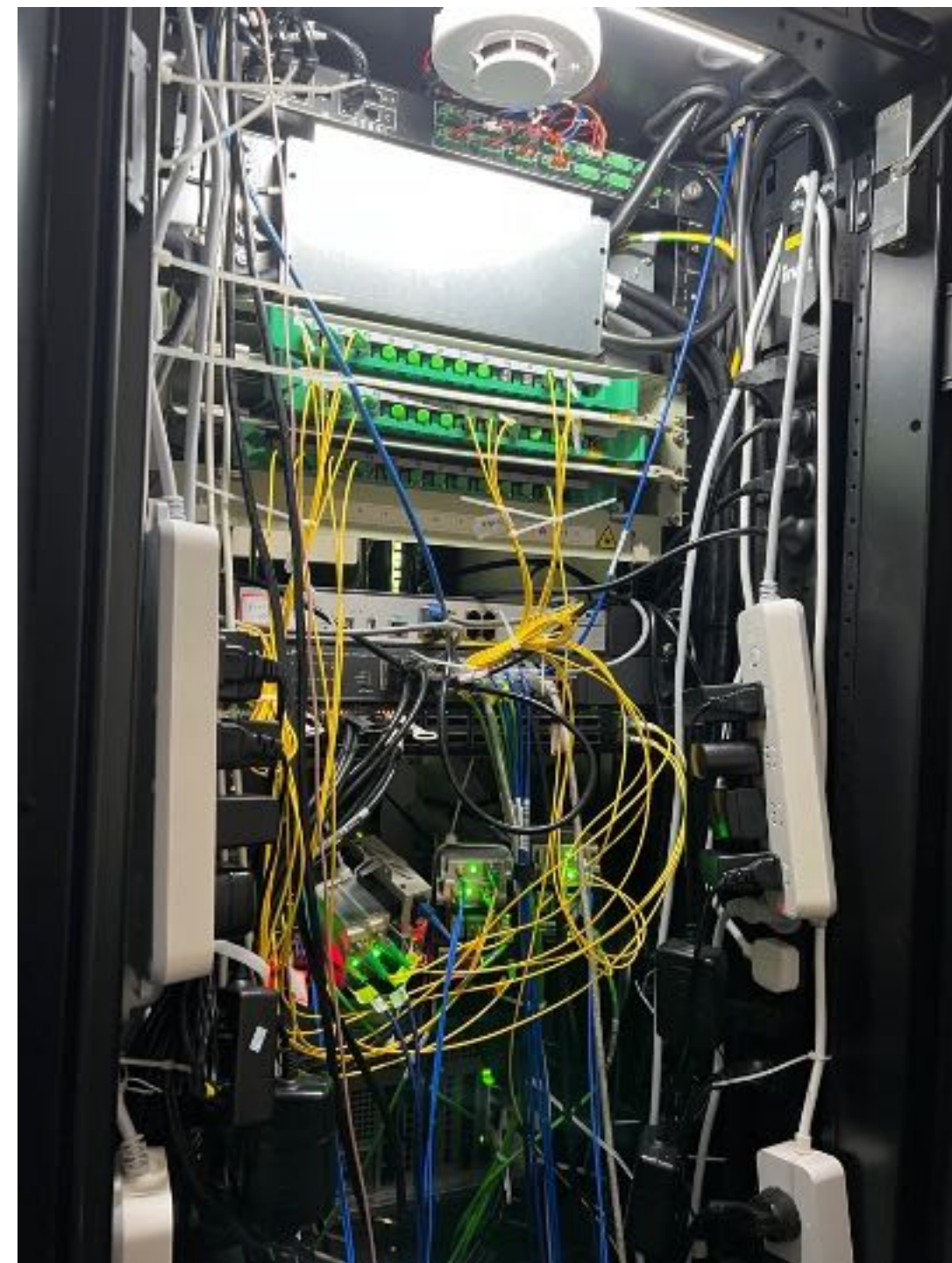
- High computing cost
- Complex algorithm
- High Data rate



Small Dishes Array in China

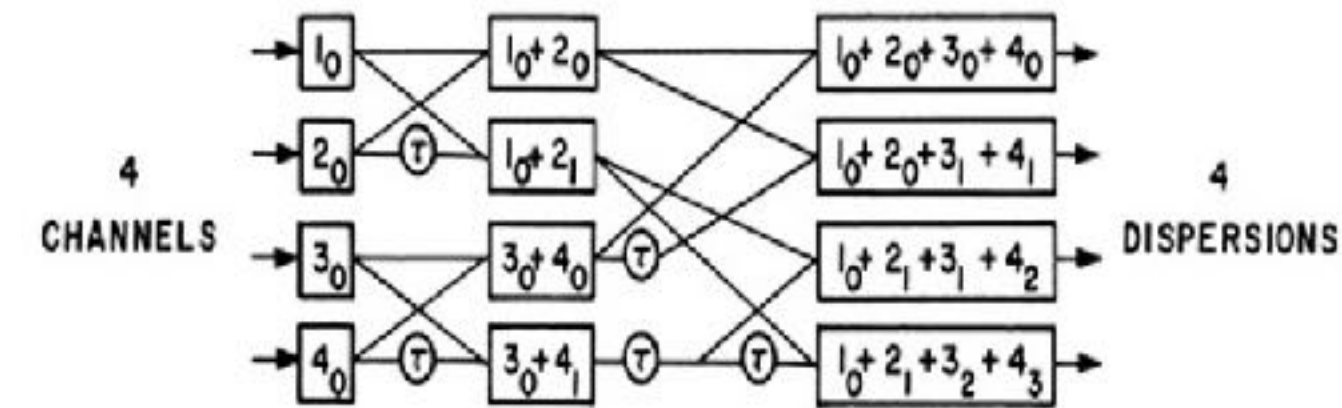
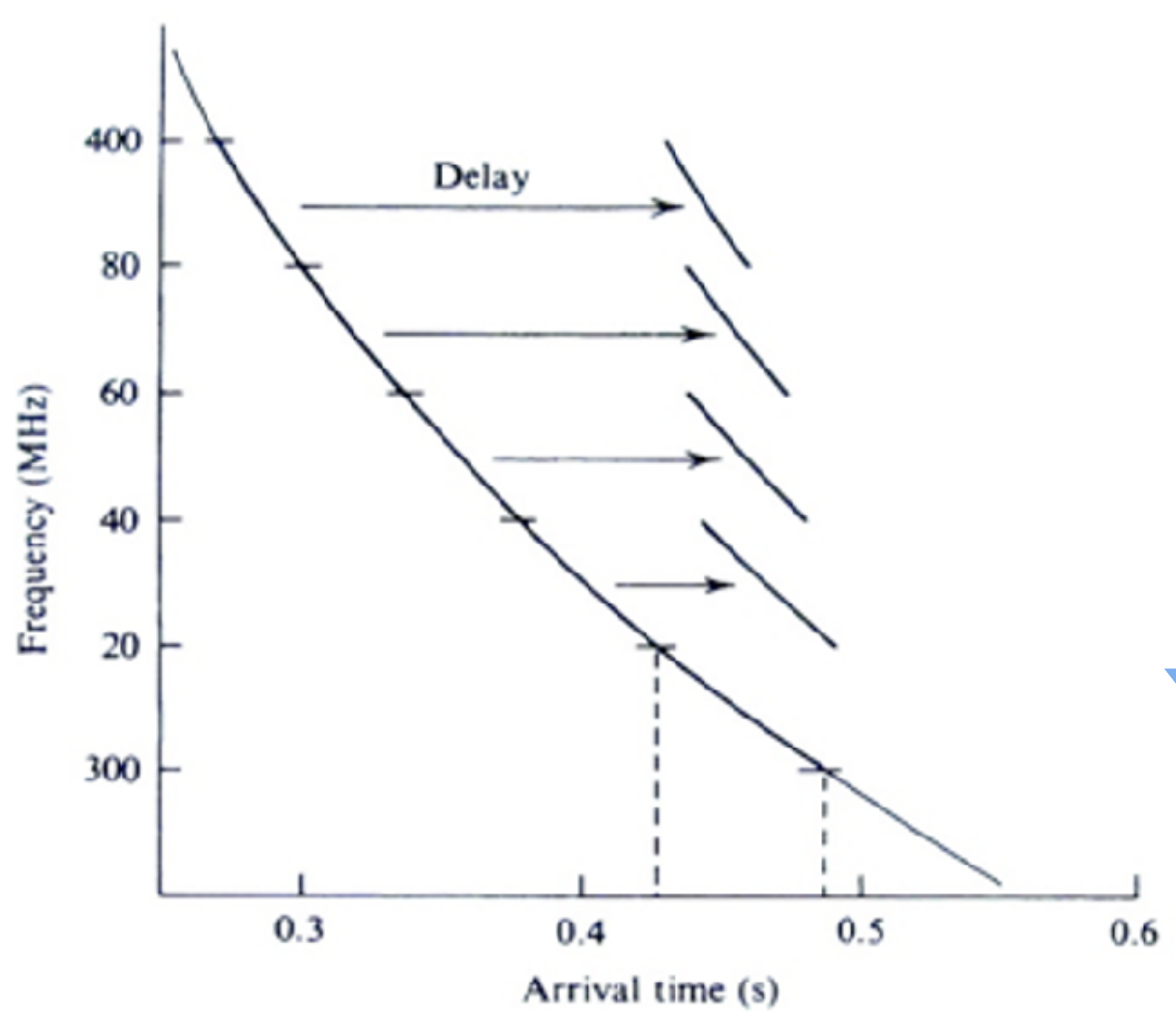
Gui Zhou Normal University
(Ru-shuang Zhao)

- 4 × 4.5 m dishes
- L-band (1-1.5 GHz)
- Cost: 200,000 Yuan per dish



How to search FRBs

De-dispersion



$$O(N_f \cdot N_t \cdot N_{DM})$$

$$O(N_f \cdot \log_2(N_f) \cdot N_t)$$

(Taylor, J. H. 1974, A&AS, 15, 367)

Time shift

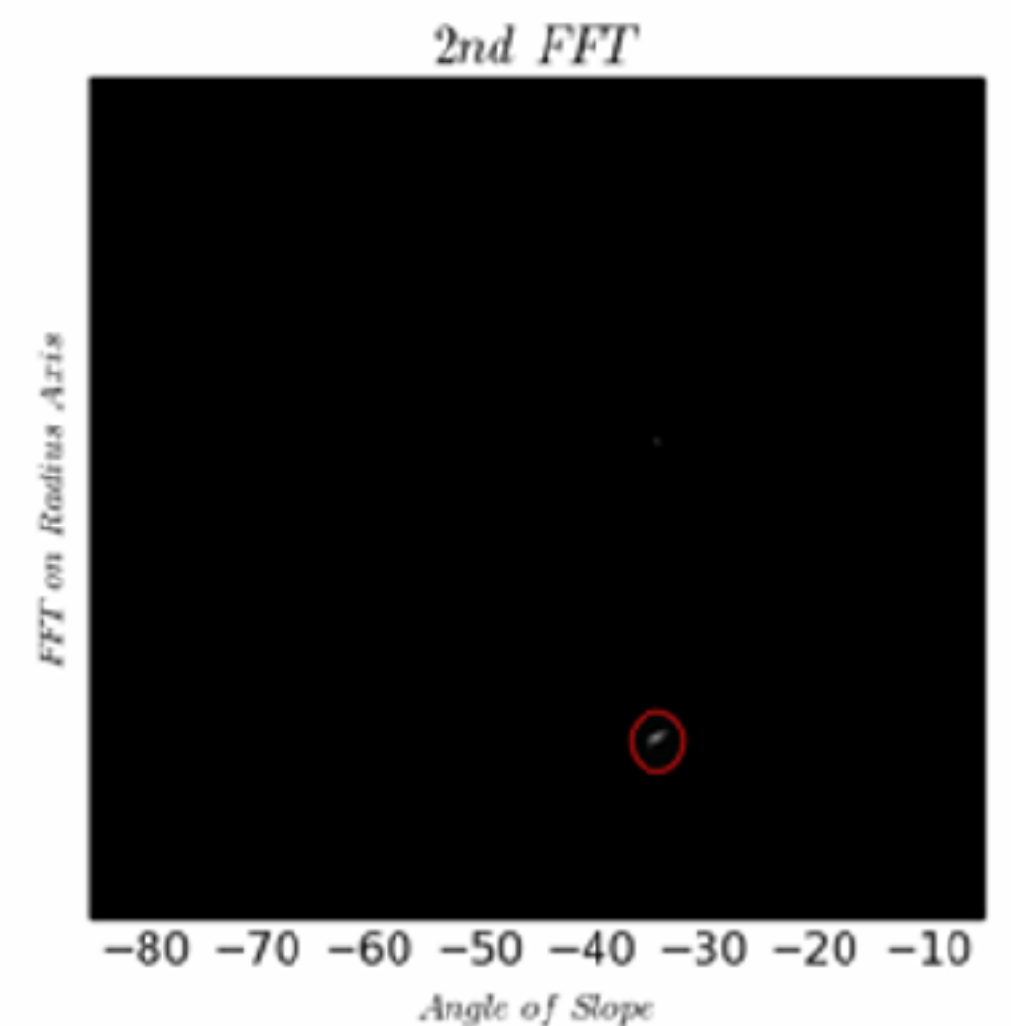
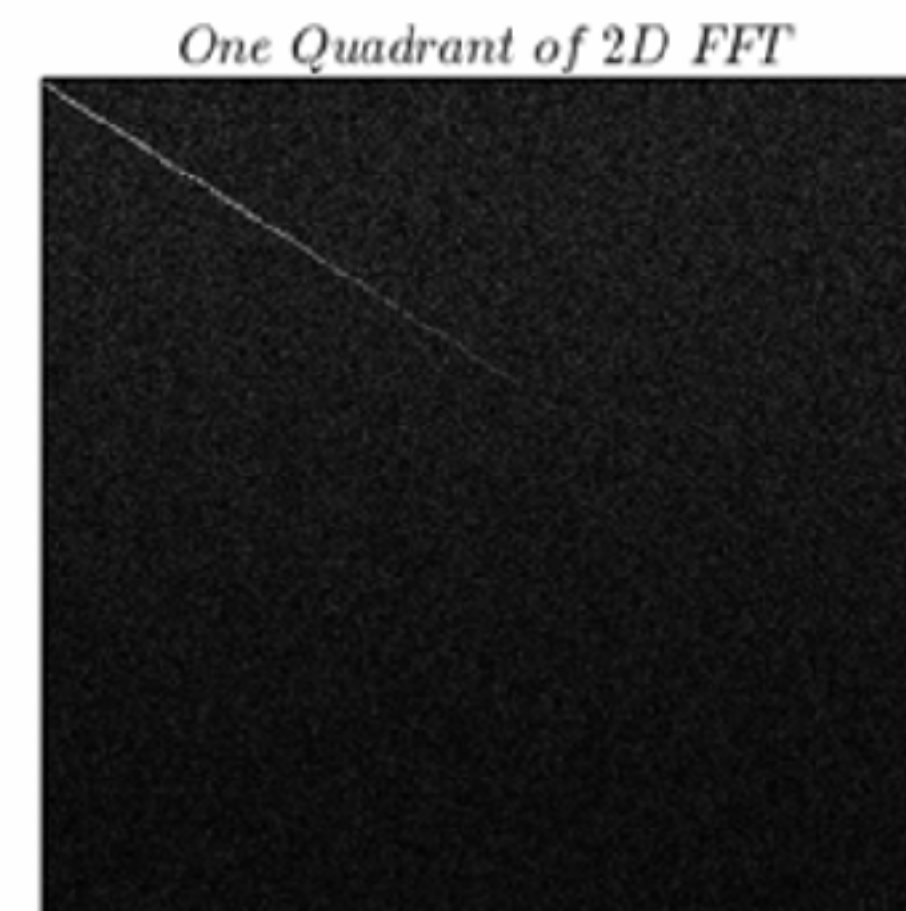
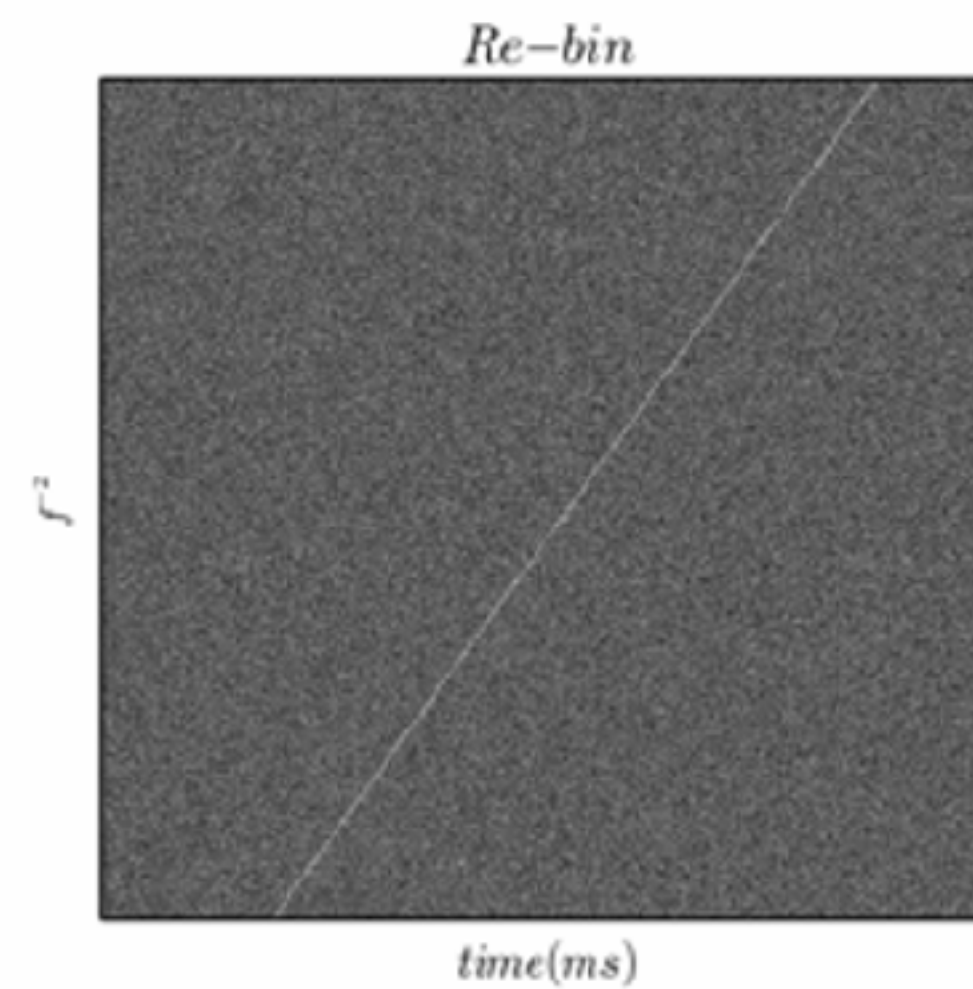
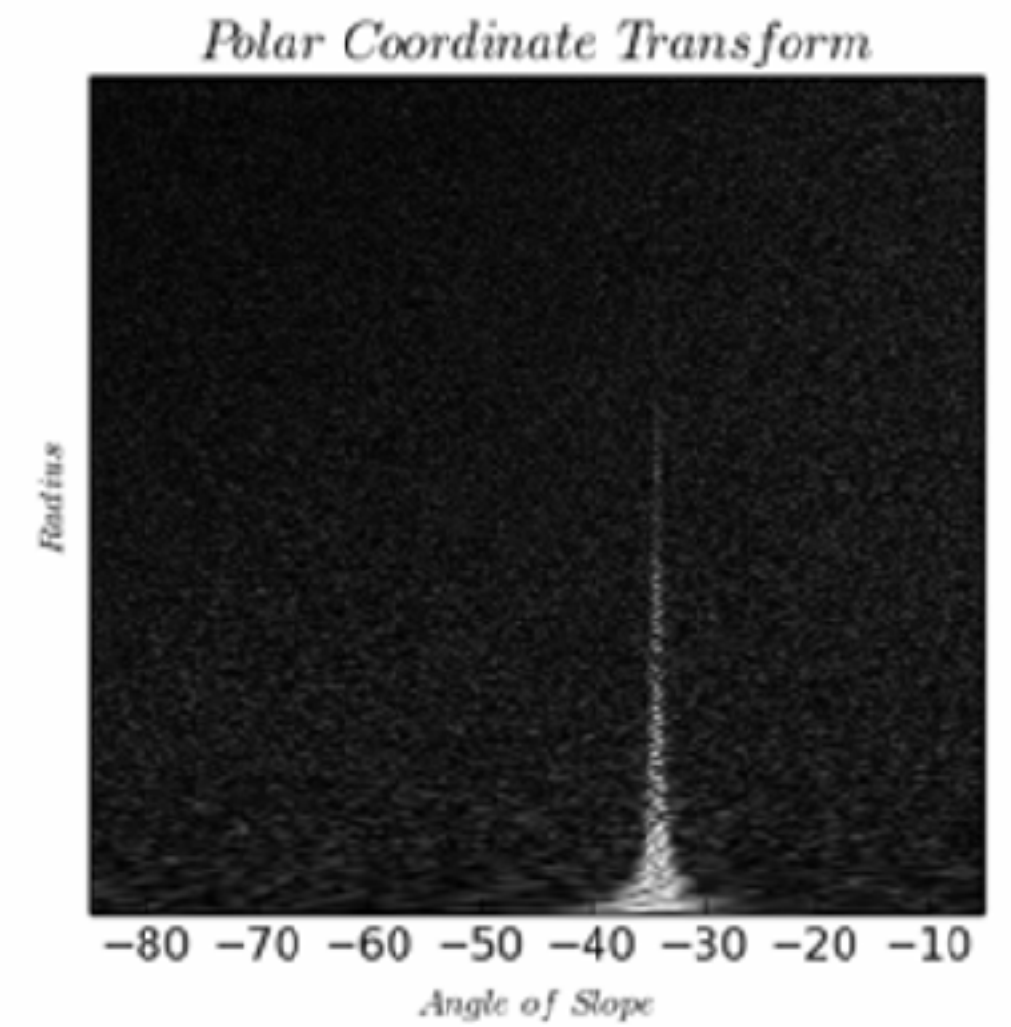
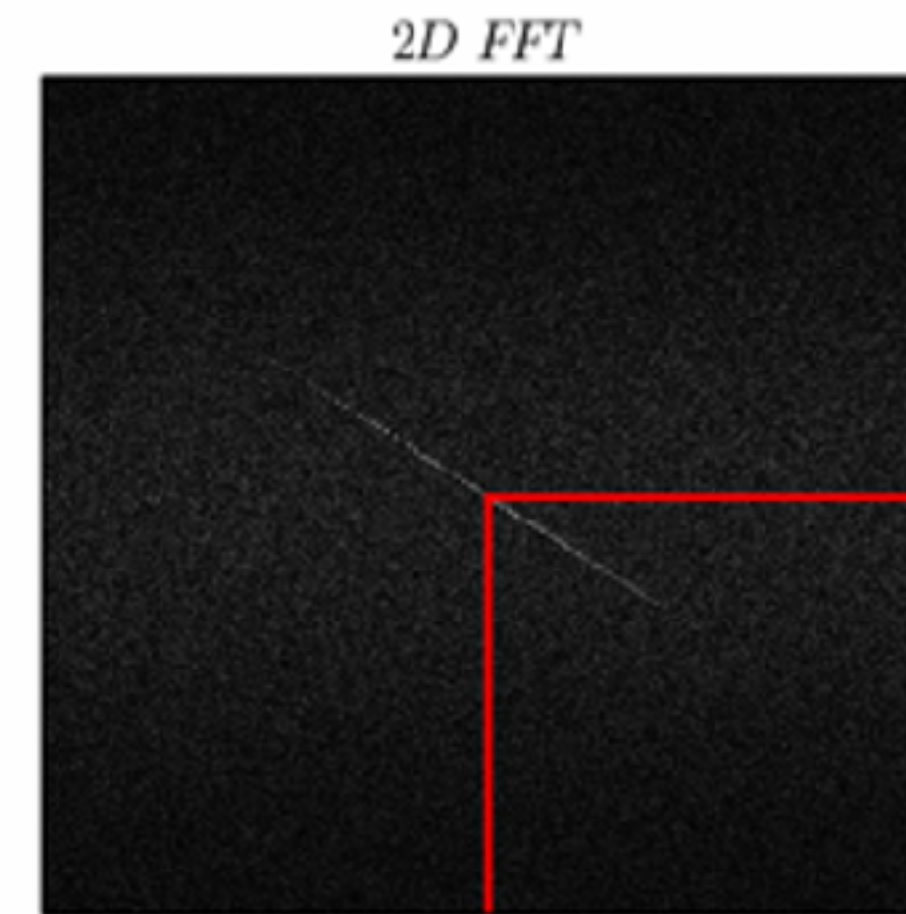
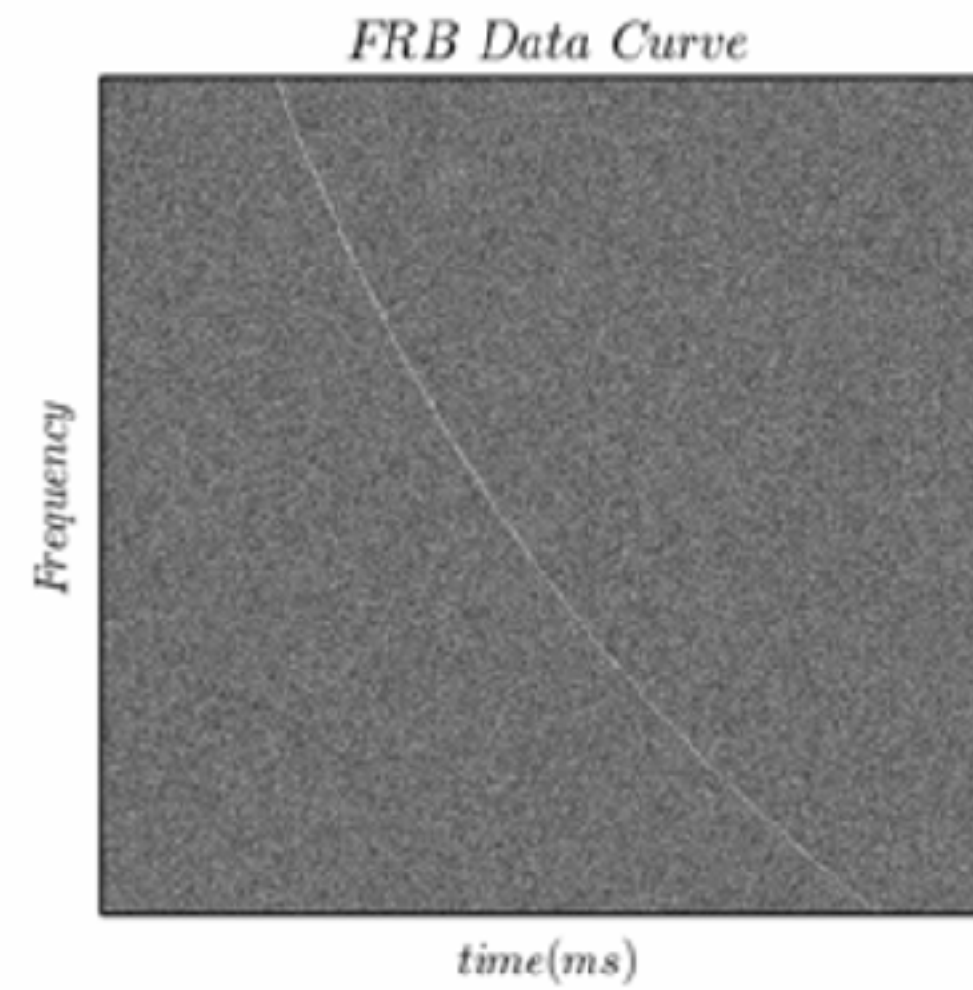
Tree Algorithm

Image Processing

Machine learning

2DFFT method

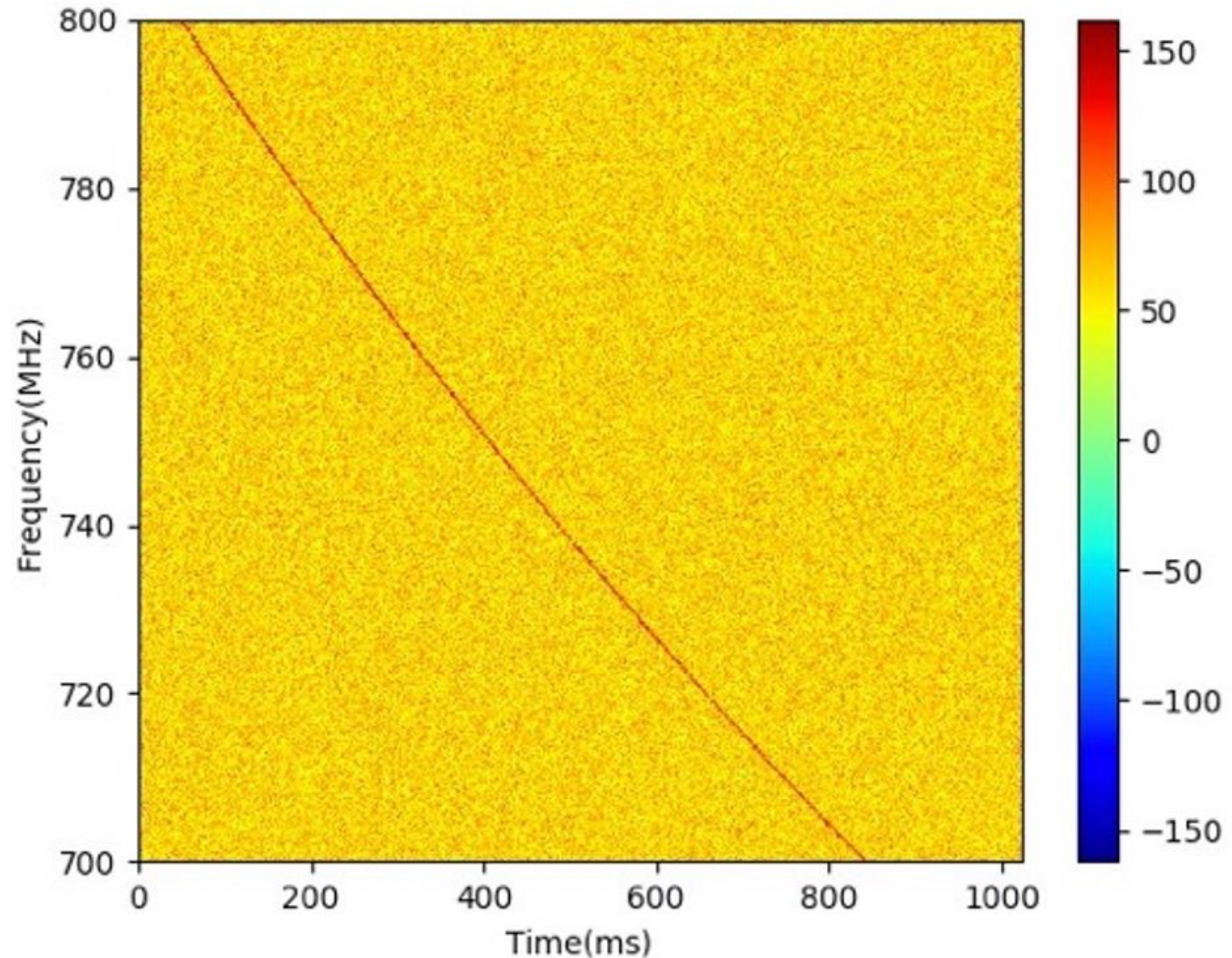
- Straighten the Curve
- 2D-FFT
- Zoom in the burst zone
- Polar Coordinates transform
- 2nd 1-D FFT along radius



$$-C \cdot DM \cdot \nu^{-2} = -t + t_0 - C \cdot DM \cdot \nu_0^{-2}$$

Simulated Burst

- **DM : 500 pc cm⁻³**
- **Time resolution : 1 ms**
- **Bandwidth : 700-800MHz**
- **Frequency Channels : 1024**
- **Total Time : 1.024s**



Straighten The Curve

Denote $\mu = \nu^{-2}$

$$\mu = \frac{1}{c \cdot \text{DM}} \cdot t - \frac{1}{c \cdot \text{DM}} \cdot t_0 + \nu_0^{-2}$$

Slope:

$$k_{re} = \frac{1}{c \cdot \text{DM}}$$

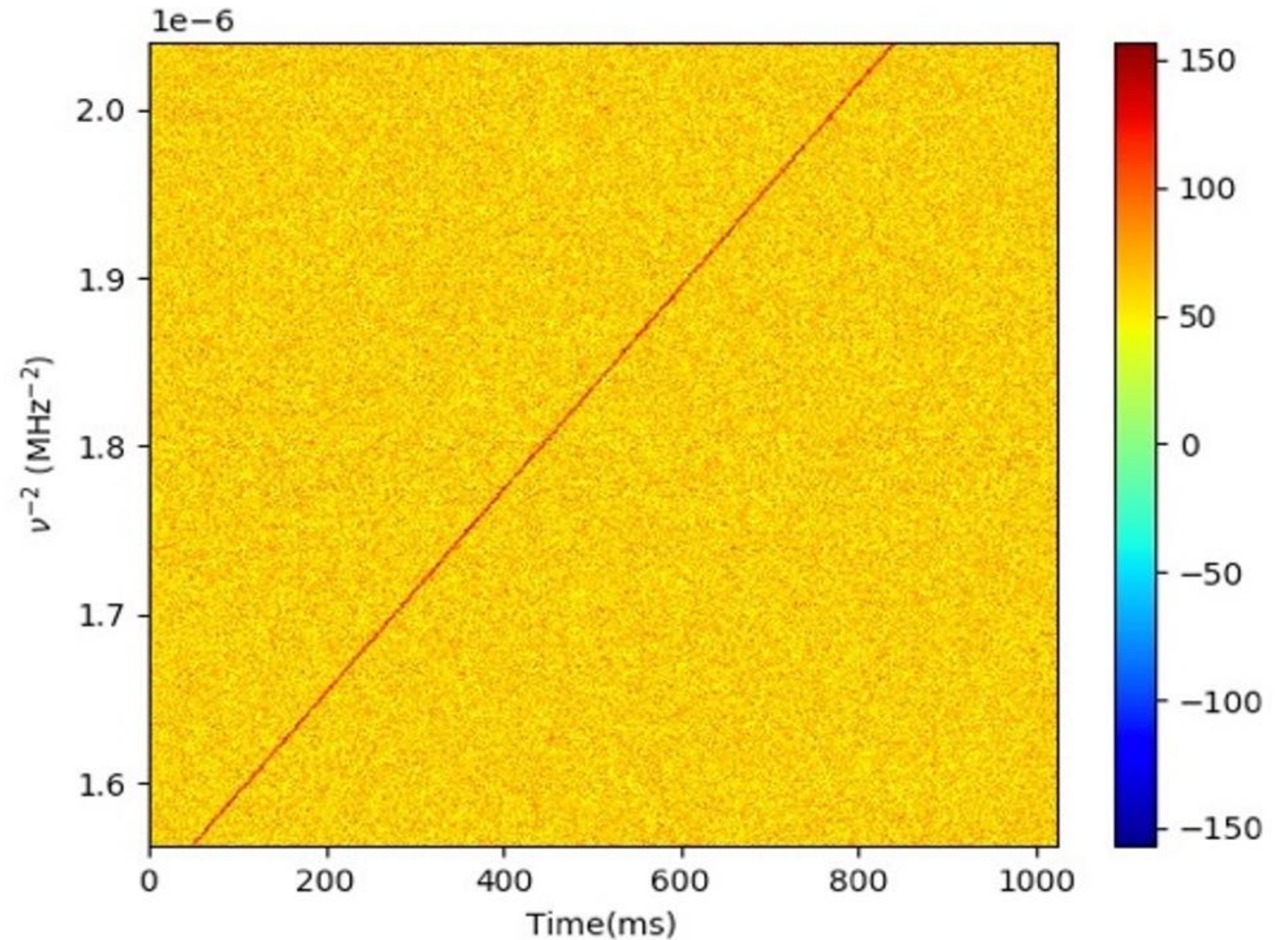
Unit of Slope:

$$\text{Unit}_k = \frac{\nu_{\min}^{-2} - \nu_{\max}^{-2}}{N_{\text{bin}} \cdot T_{\text{samp}}}$$

Interception of vertical axis:

$$b = -\frac{1}{c \cdot \text{DM}} \cdot t_0 + \nu_0^{-2}$$

Where $t_0 = 50$ ms



2D FFT on the straight

In the re-bin map, the curve map could be described as:

$$I(\mu, t) \cdot \delta(k_{rebin} \cdot t + b - \mu)$$

Take Fourier Transfer pair: $(\mu, t) \sim (\omega, f)$, then take 2DFFT:

$$\begin{aligned} & \mathcal{F}_{2D} \{ I(\mu, t) \cdot \delta(k_{re}t + b - \mu) \} \\ &= \iint I(\mu, t) \cdot \delta(k_{re}t + b - \mu) \cdot e^{-i2\pi(\omega\mu+ft)} d\mu dt \\ &= \int I(k_{re}t + b, t) \cdot e^{-i2\pi(\omega(k_{re}t+b)+ft)} dt \\ &= \int I(k_{re}t + b, t) \cdot e^{-i2\pi(\omega k_{re}+f)t} \cdot e^{-i2\pi\omega b} dt \\ &= I_{total} \cdot \delta(\omega k_{re} + f) \cdot e^{-i2\pi\omega b} \end{aligned}$$

Transient in (ω, f) plot

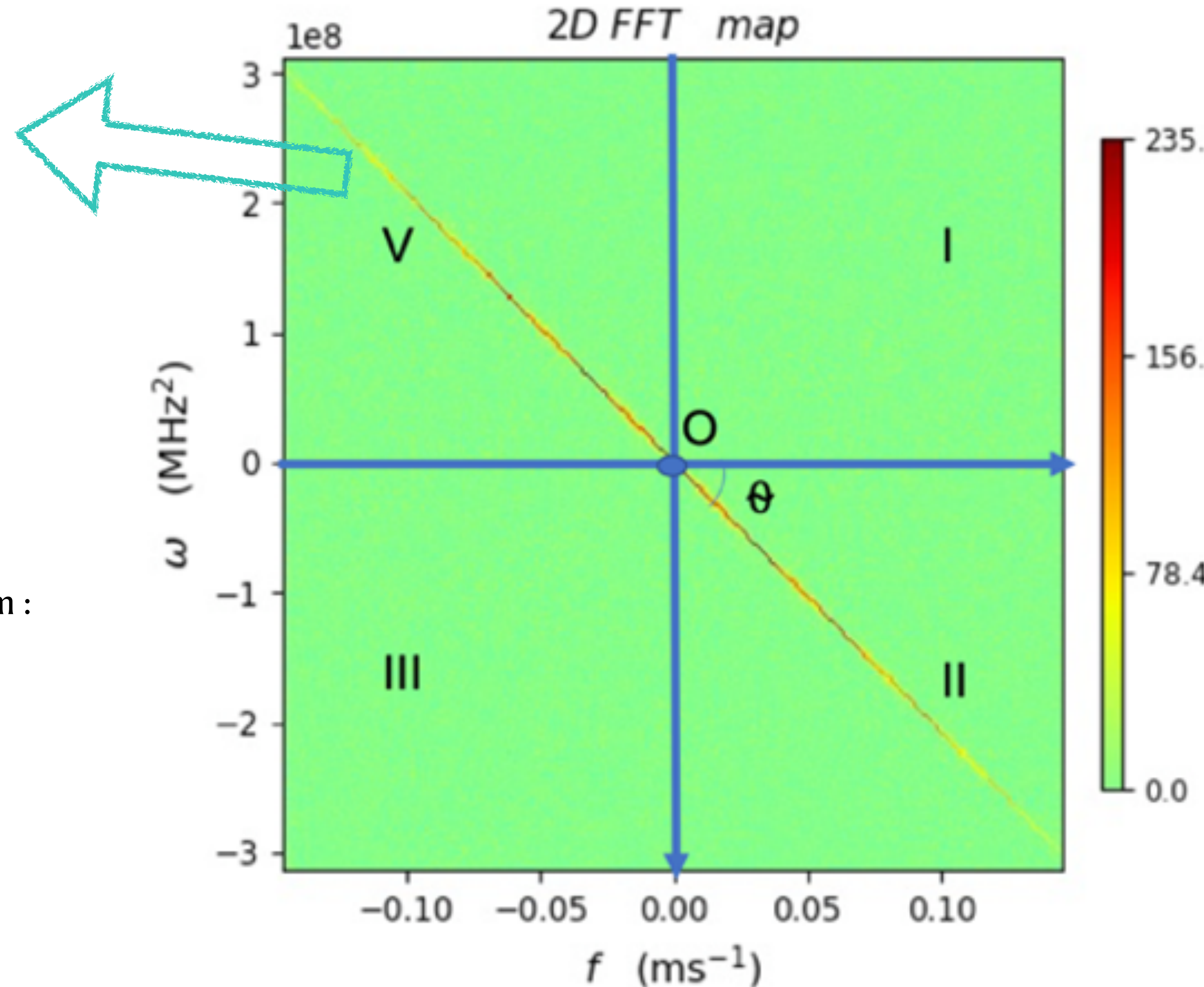
$$I_{total} \cdot \delta(\omega k_{re} + f) \cdot e^{-i2\pi\omega b}$$

In $I(\omega, f)$ plot, The slope of straight is:

$$\omega = -\frac{1}{k_{re}} f$$

And the interception in (μ, t) plot goes to the phase term :

$$e^{-i2\pi\omega b}$$



Transient in (θ, r) plot

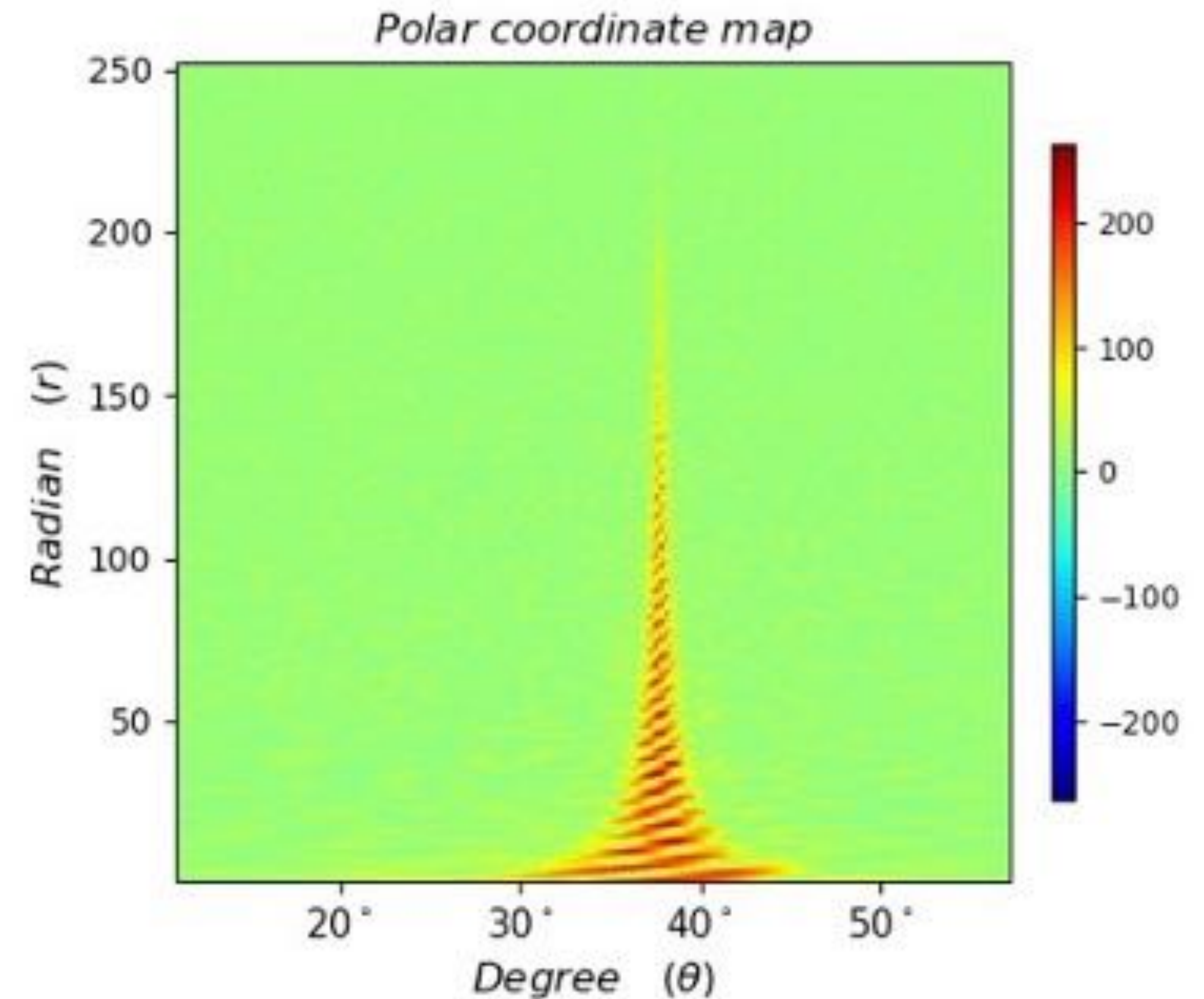
$$\begin{cases} f = r \cdot \cos \theta \\ \omega = N_\omega - r \cdot \sin \theta \end{cases}$$

The standard deviation of 2D matrix of (θ, r) map:

$$\sigma_{2D} = \sqrt{\frac{\sum_{i=1}^{N_\theta} \sum_{j=1}^{N_r} (I'_{ij} - \bar{I}')^2}{(N_\theta - 1)(N_r - 1)}}$$

Sum along r axis:

$$\sigma_i = \sqrt{\frac{\sum_{i=1}^{N_r} (I'_i - \bar{I}')^2}{(N_\theta - 1)}}$$

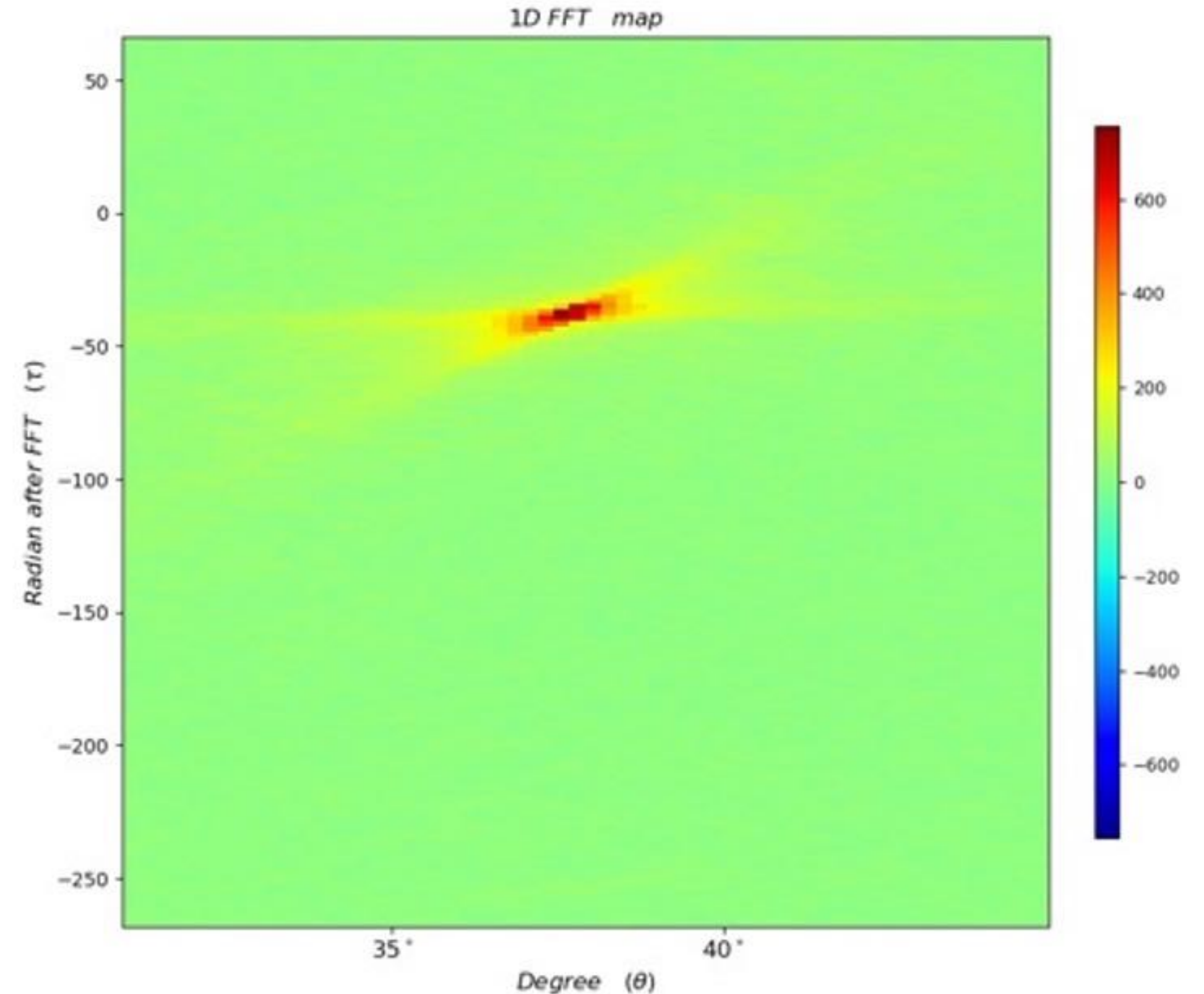


The significance after additional 1D FFT

Second 1-D FFT along radian direction, Then the signal converge into some dots.

The amplitude square value of matrix elements contribute to χ^2 distribution with degree 2. Pick the singular values from the spots. Then calculate the significance as:

$$r_{sig} = \sum_{\Delta\theta} \frac{|I(\tau, \theta)|_{\geq T_\sigma}^2}{\sigma \cdot N_{pixel}}$$



2DFFT Pipeline

Based on this algorithm we developed a open source pipeline which called 2DFFT:
(https://github.com/peterniuzai/2DFFT-transient_search_Pipeline)

```
python __main__.py -h:
```

```
Usage: mpirun -n <number of processor> python __main__.py [options]
```

```
Options:
```

```
-h, --help show this help message and exit
```

```
-f FILE, --file=FILE Put filterbank file want to search
```

```
-t THRESHOLD, --threshold=THRESHOLD
```

```
                Threshold(sigma) for candidates pick
```

```
--dm=DM          Set DM range, Suggest use default [50,2000]
```

```
--pixel=PIXEL    Pixels number at 2nd 1DFFT map
```

```
--mask_cycle=MSK_CYCLE
```

```
                Mask abnormal lines at 1st 2DFFT map, only when RFI is terrible
```

```
--nsamps_gulp=T_LEN Samples number for onece process, suggest use self-calculate(def)
```

```
--nbin=NBIN      number of channels after re-bin step, suggest use self-calculate(def)
```

```
--wp=WP          Set width of pulse to search, Suggest use default 10(ms).
```

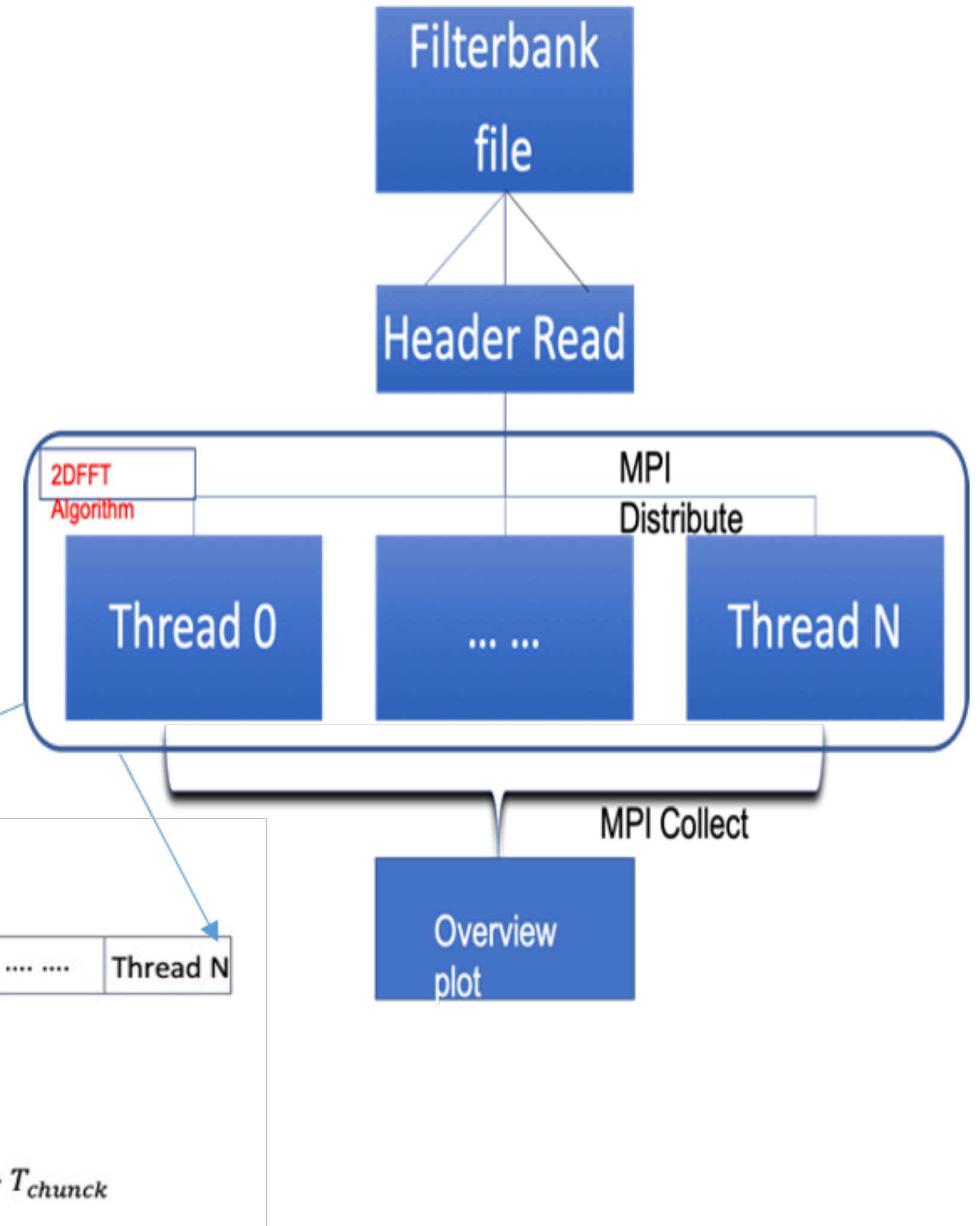
```
--angle=ANGLE    Angle range within [0,90], For de-buging.
```

```
-v, --verbose    Show details of process
```

```
-p, --plot       Make overview plot for final result
```

```
--Plot_proc=PLOT_PROC
```

```
                Input process step key words want to make single plot. Key words Including: {raw, rebin, 1stFFT, polar_sets_3D, polar_sets_2D, 2ndFFT_3D, 2ndFFT_2D} (This function Remain updates)
```



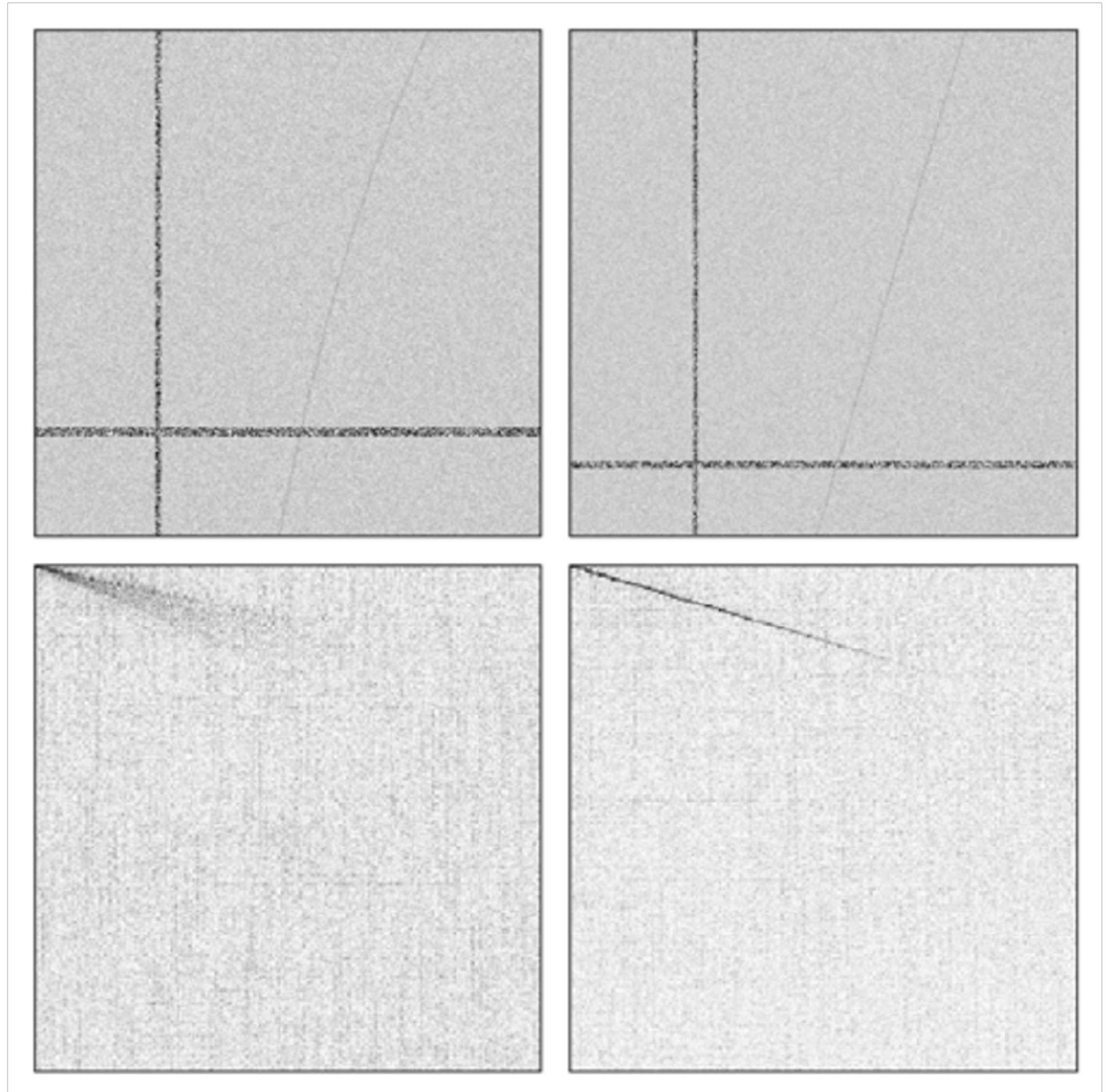
Better at large scale RFI eliminating

Top panels: Left and right plots the transient waterfall with and without pre-process.

Bottom panels: Perform 2D FFT on the transients with and without the pre-process.

Results:

- After 2DFFT, The transient divergence with pre-process is much smaller than that without pre-process.
- The large scale RFI could be eliminated clearly after 2DFFT.



The S/N improvement after re-bin process

Time resolution : 1 ms

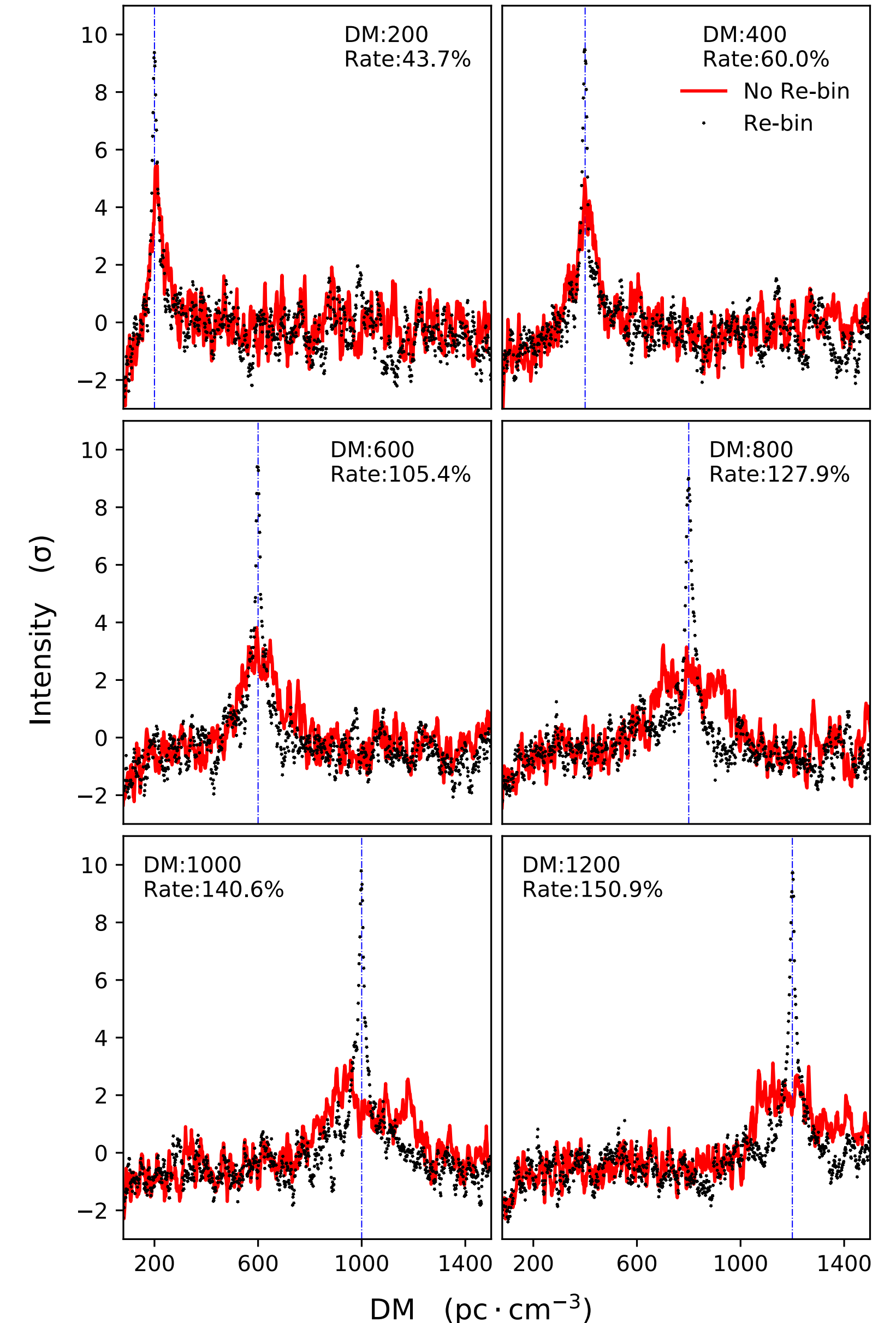
Frequency Bandwidth: 700-800 MHz

Number of frequency Channels : 4096

$$S = (I_{max} - I_{mean}) / \sigma$$
$$\mathcal{R} = (S_{rebin} - S_o) / S_o$$

Results:

- The S/N improvement Varied with DM. (Curvature)
- Due to the large DM of FRB, Re-bin process is necessary to be taken in use.



The detection rate after additional 1D FFT

Time resolution : 1 ms

Frequency Bandwidth: 700-800 MHz

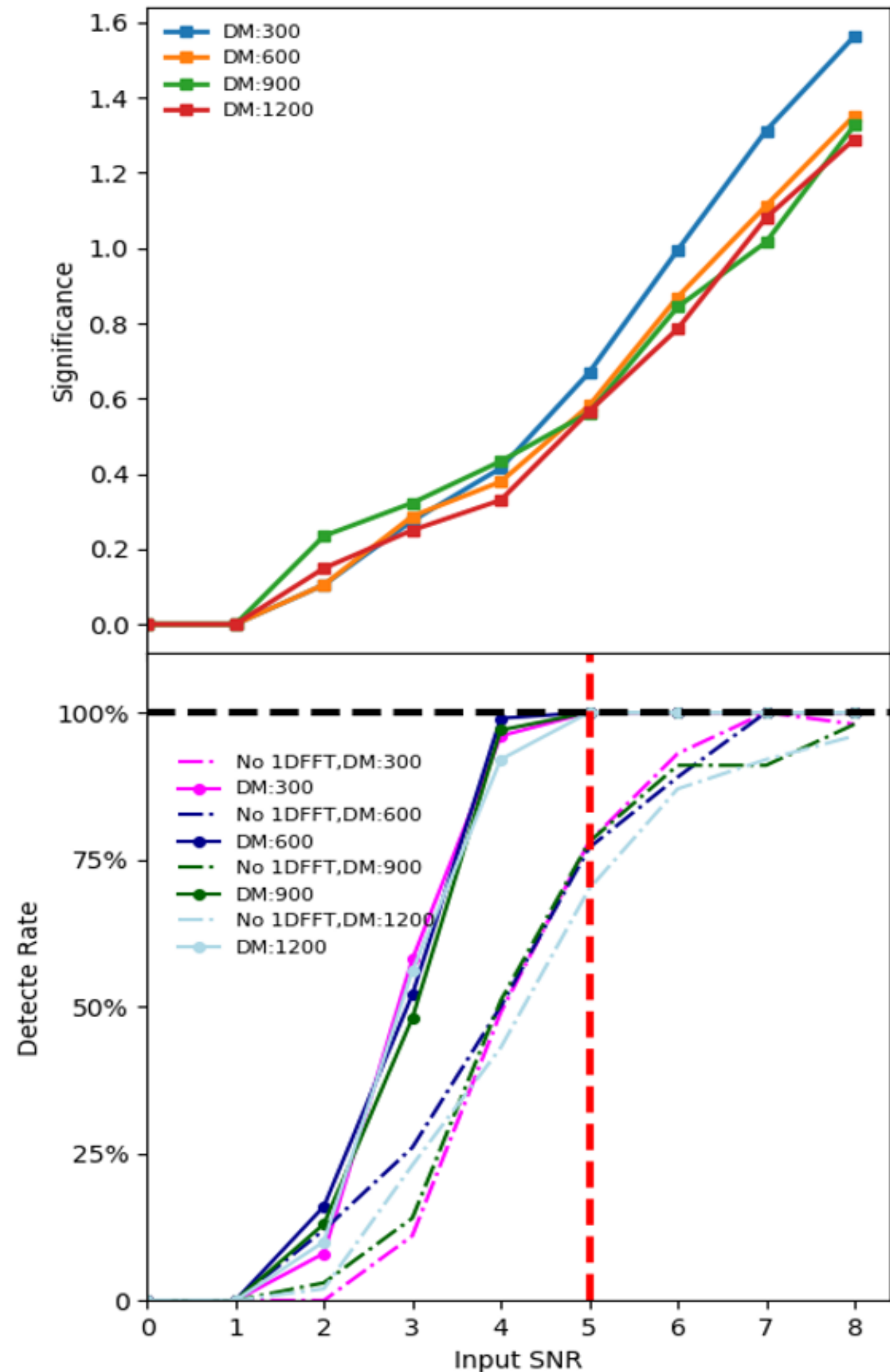
Number of frequency Channels : 1024

Inject different DM transients

Each DM are generated 100 times.

Results:

- Result: With additional 1DFFT, the detection sensitivity could reach 100% when input SNR is 5.
- Additional 1D FFT is necessary and helpful to improve the detection rate.



Time cost of each step

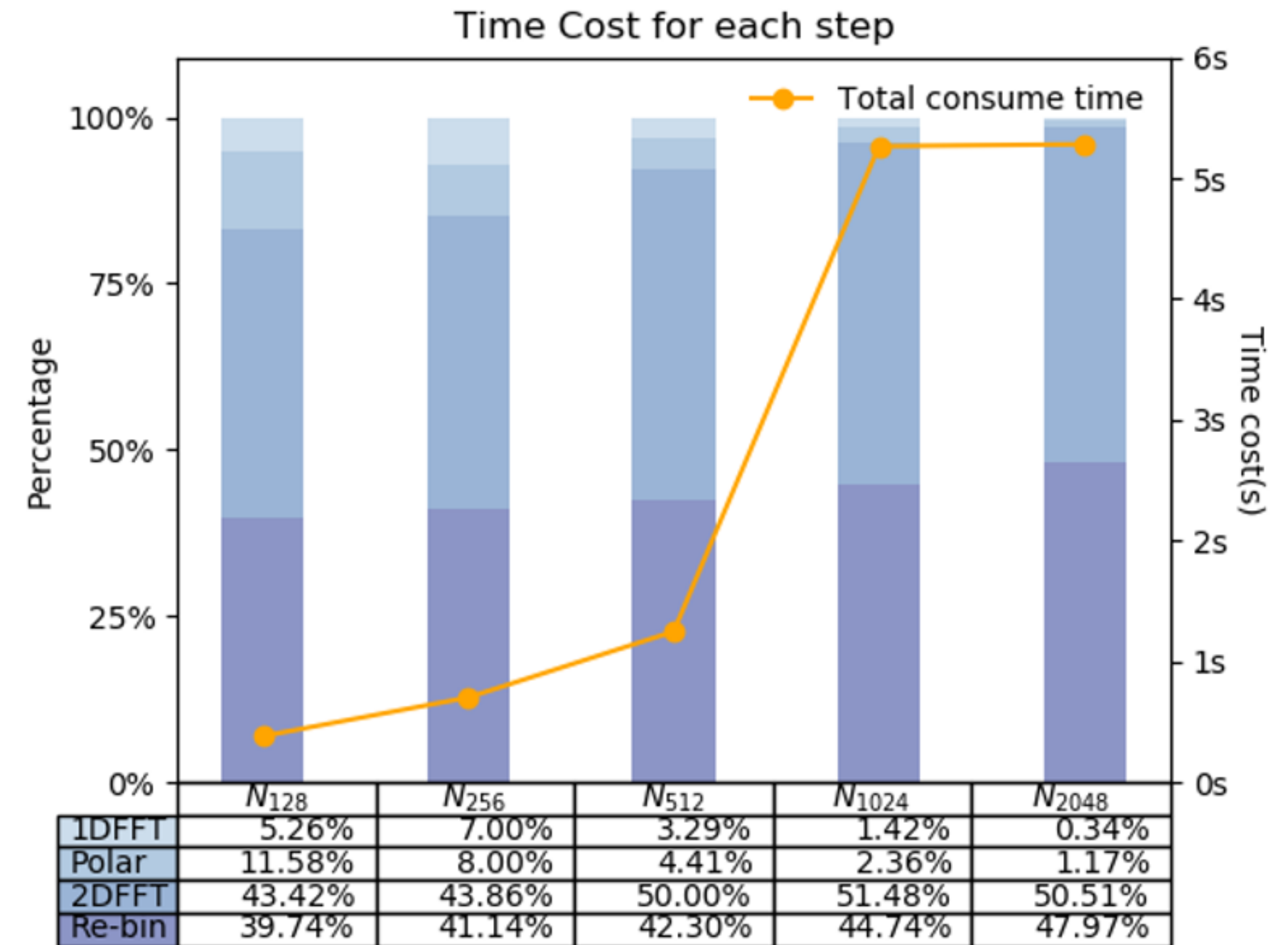
Time resolution : 1 ms

Frequency Bandwidth: 700-800 MHz

Number of frequency Channels : ?

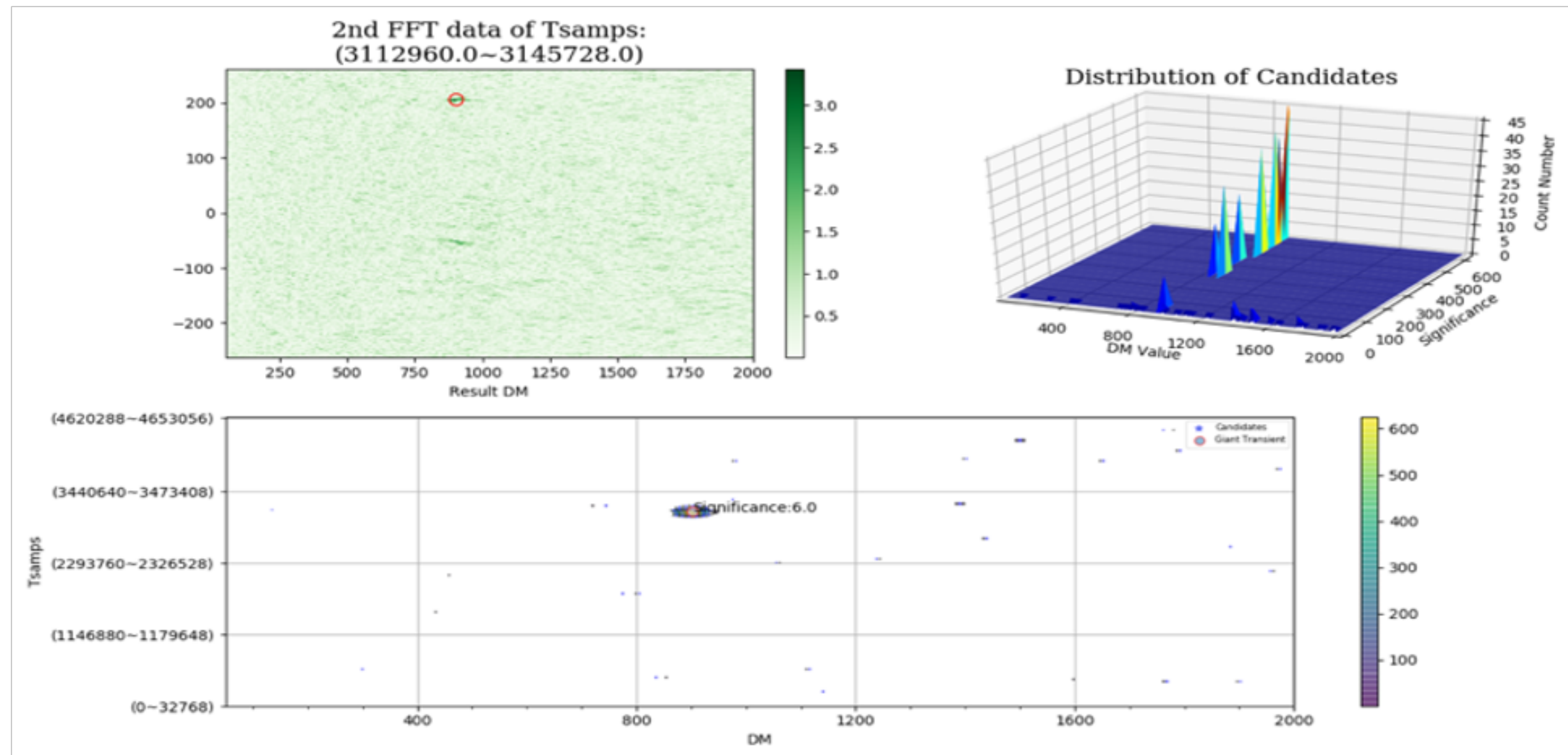
Results:

- 2DFFT and re-bin step cost a lot in this algorithm
- The other steps are acceptable



Exam the known FRBs

FRB name	$n_t \times n_f$	Data Duration(m)	Process time(m)	SNR	Width(ms)	DM	DM Found
FRB140514	14438400 × 1024	15.4	6.8	16	2.8	562.7±0.6	567.96
FRB010125	2121728 × 96	4.4	5.5(s)	17	9.4	790±3	799.21
FRB010621	8400896 × 96	35.0	0.38	16.3	7	745±10	747.89
FRB090625	4620288 × 1024	4.9	2.23	30	1.92	899.55±0.01	897.03
FRB130729	4554752 × 1024	4.9	2.1	14	15.61	861±2	860.46
FRB130628	4587520 × 1024	4.9	2.2	29	0.64	469.88±0.01	470.46
FRB130626	4587520 × 1024	4.9	2.1	21	1.98	952.4±0.1	951.88
FRB110220	4489216 × 1024	4.8	2.1	49	5.6	944.38±0.05	945.78
FRB110523	55296 × 4096	56.62(s)	5.18(s)	42	1.73	623.3±0.06	627.26



Compare with Other Open Source Software

Time resolution : 1 ms

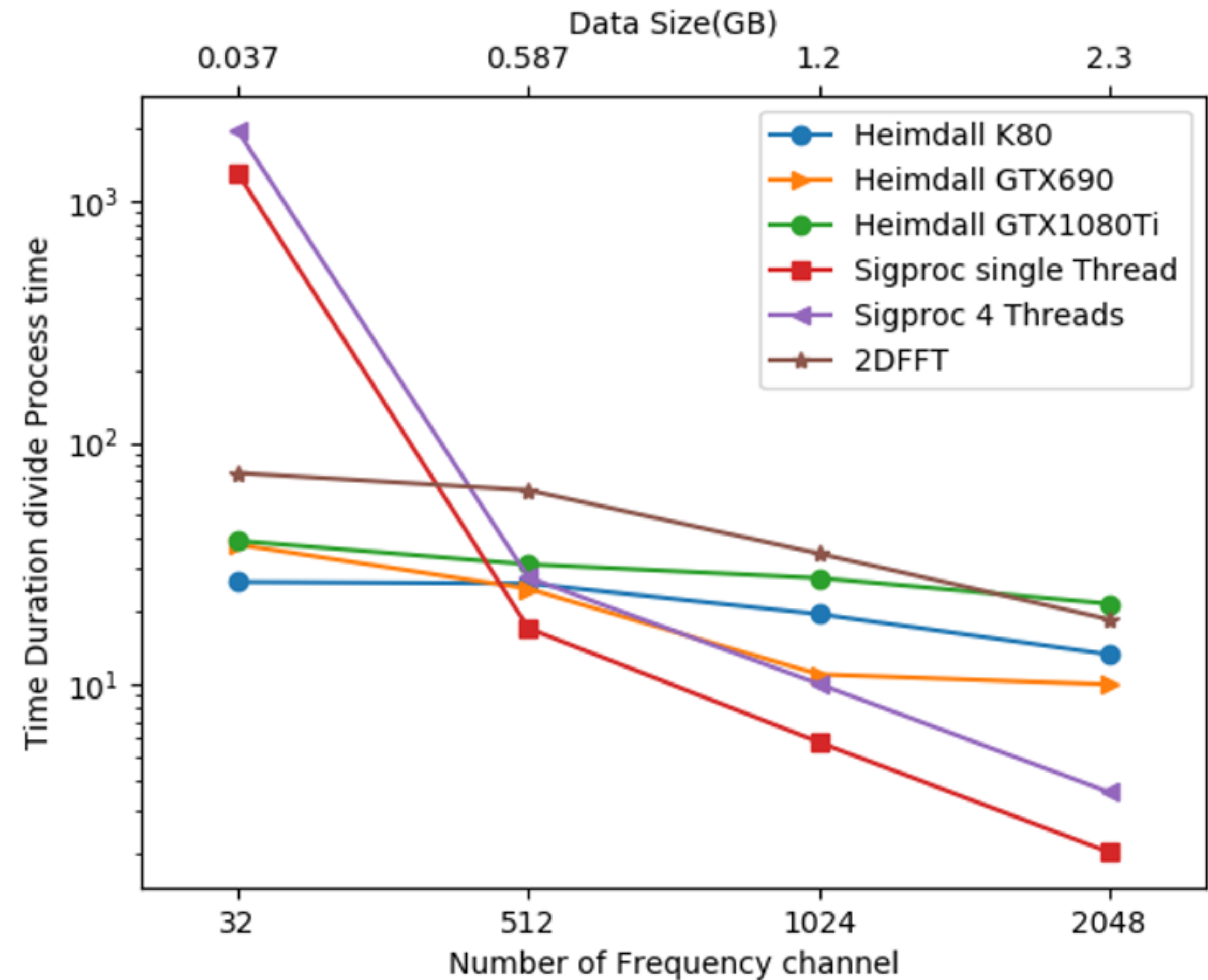
Frequency Bandwidth: 700-800 MHz

Number of frequency Channels : -?

Total Inject time duration: 1200 s

Software list:

- Heimdall (GPU based)
- SIGPROC (Single and multi Treads)
- SIGPYPROC
- 2DFFT (4 Threads)



Compare with Other Open Source Software

Time resolution : 1 ms

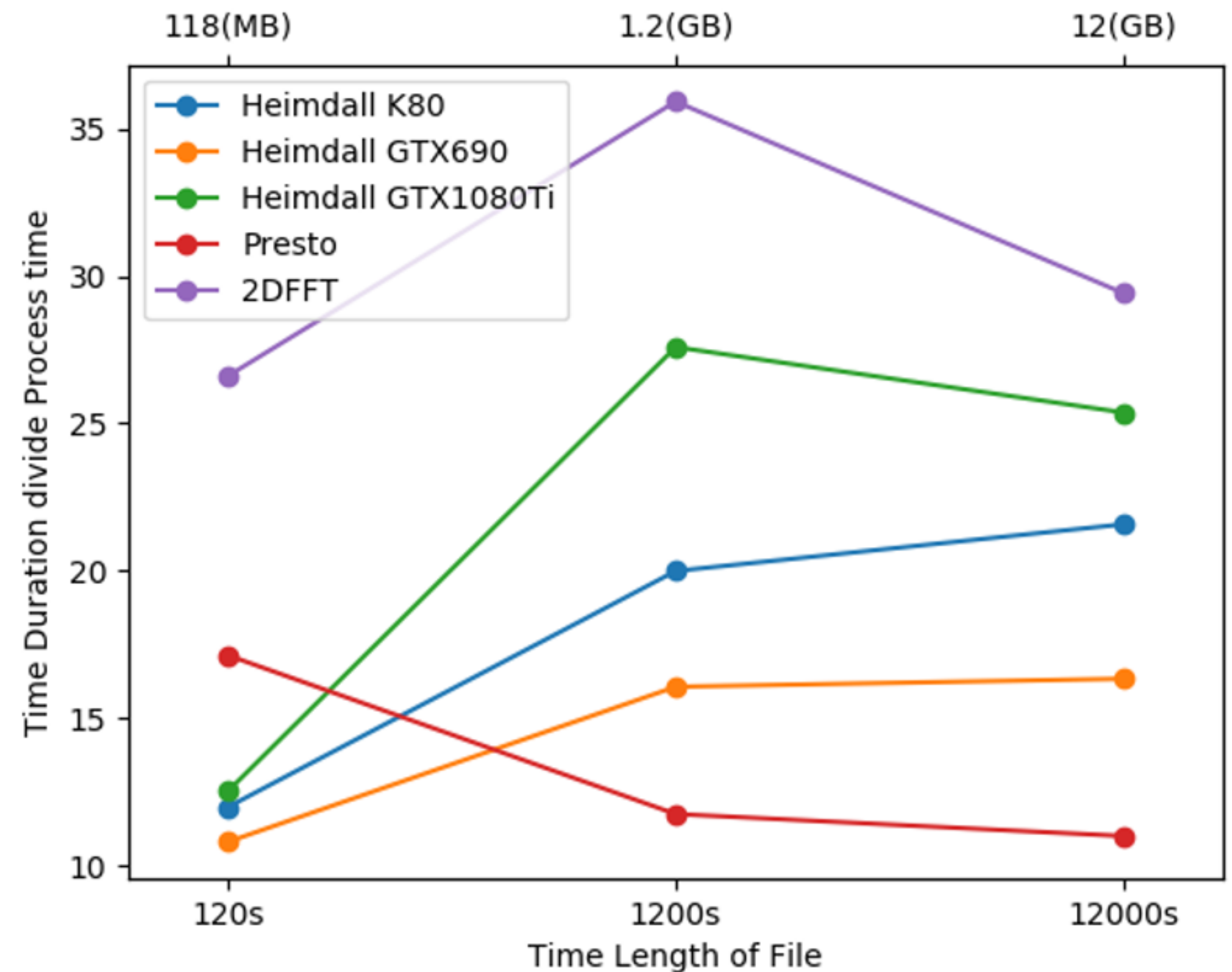
Frequency Bandwidth: 700-800 MHz

Number of frequency Channels : 1024

Total Inject time duration: -?

Software list:

- Heimdall(GPU based)
- PRESTO (Sub-band)
- 2DFFT(4 Threads)



Compare with Other Open Source Software

Time resolution : 0.1 ms

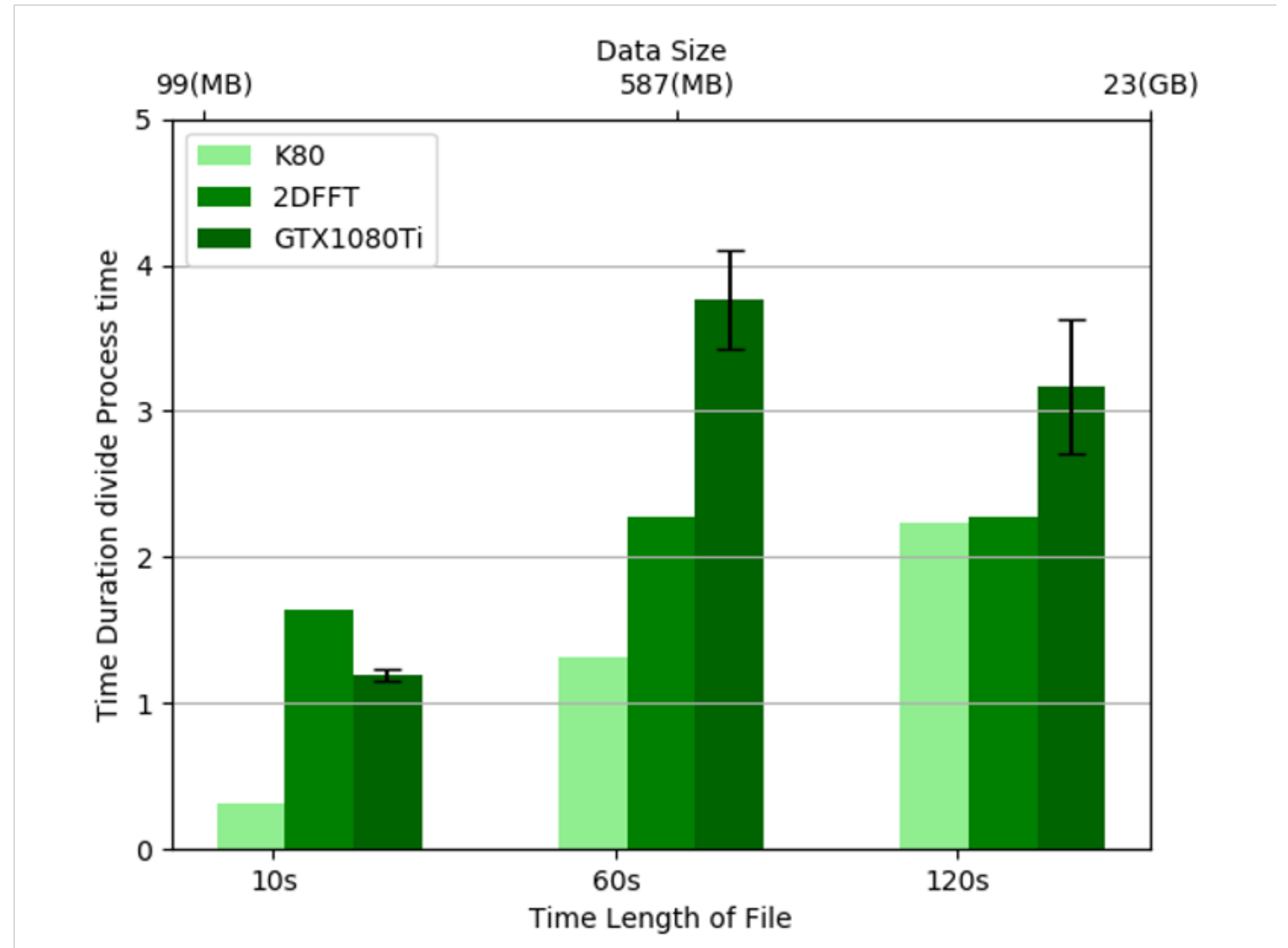
Frequency Bandwidth: 700-800 MHz

Number of frequency Channels : 2048

Total Inject time duration: -?

Software list:

- Heimdall(GPU based)
- 2DFFT(4 Threads)



Conclusion & Future

2DFFT is a highly effective pipeline for searching radio transients. It excels at eliminating radio frequency interference (RFI) and parallel processing. Although the time cost of the re-bin step could not be ignored, the pipeline based on GPU is also promising.

Our collaborators from Computer Network Information Center (CNIC) have created a GPU-accelerated 2DFFT pipeline. This pipeline has the capability to achieve twice the speed of our Python version. More information regarding the specifics will be provided soon.